# Kinetic theory of gases I

The Gibbs statistical mechanics is applicable to systems in thermodynamical equilibrium. In the subsequent lectures I will presented the second pillar of statistical mechanics – kinetic theory which can be applied to both equilibrium and non-equilibrium systems. The kinetic theory also aspires to explain unidirectional temporal evolution towards thermodynamical equilibrium.

The kinetic theory initiated by James Clerk Maxwell (1831-1879) and further developed by Ludwig Boltzmann (1844-1906) is a rich theory of various applications. My presentation, however, will be limited to <u>dilute</u> gases where atomic sizes are much smaller than inter-particle distances. Consequently gas particles are mostly non-interacting and there are interactions (collisions) of no more than two particles.

# **Distribution function**

• The main object of kinetic theory is the <u>distribution function</u>  $f(t, \mathbf{r}, \mathbf{p})$  where  $\mathbf{r}$  and  $\mathbf{p}$  is the particle position and momentum and t denotes the time. The quantity

$$f(t, \mathbf{r}, \mathbf{p}) \frac{d^3 r \, d^3 p}{(2\pi)^3} \tag{1}$$

tell us how many particles there are in the phase-space volume  $\frac{d^3r d^3p}{(2\pi)^3}$  around  $(\mathbf{r}, \mathbf{p})$  at time t. The volume is infinitesimally small from a macroscopic point of view but it is microscopically so big that there are many gas particles in the volume. We note that there are  $10^{19}$  atoms in cm<sup>3</sup> of the air in normal conditions.

• The integral

$$\int \frac{d^3 r \, d^3 p}{(2\pi)^3} f(t, \mathbf{r}, \mathbf{p}) = N \tag{2}$$

gives the total number of gas particles.

• The time-dependent particle density and flux are

$$\int \frac{d^3p}{(2\pi)^3} f(t, \mathbf{r}, \mathbf{p}) = \rho(t, \mathbf{r}), \qquad (3)$$

$$\int \frac{d^3 p}{(2\pi)^3} \, \frac{\mathbf{p}}{m} \, f(t, \mathbf{r}, \mathbf{p}) = \mathbf{j}(t, \mathbf{r}), \tag{4}$$

where m and  $\mathbf{v} \equiv \frac{\mathbf{p}}{m}$  are the particle mass and velocity, respectively.

• Since the gas is assumed to be dilute the energy of gas particles is mostly kinetic  $\epsilon_{\mathbf{p}} = \frac{\mathbf{p}^2}{2m}$  and the gas energy density equals

$$\int \frac{d^3p}{(2\pi)^3} \epsilon_{\mathbf{p}} f(t, \mathbf{r}, \mathbf{p}) = \varepsilon(t, \mathbf{r}).$$
(5)

• If the gas is in the external field of potential forces, *e.g.* gravitational, and the particle potential energy is  $v(\mathbf{r})$ , the gas energy density is

$$\int \frac{d^3p}{(2\pi)^3} \left(\frac{\mathbf{p}^2}{2m} + v(\mathbf{r})\right) f(t, \mathbf{r}, \mathbf{p}) = \varepsilon(t, \mathbf{r}).$$
(6)

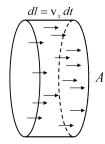


Figure 1: The pressure exerted on the surface A results from the particle collisions with the surface.

### Gas in thermal equilibrium

Predictions of kinetic theory obviously agree with those of Gibbs mechanics in case of thermodynamical equilibrium. Nevertheless I will derive two simple already known formulas to show how the kinetic theory works.

• Keeping in mind the partition function of ideal gas, one guess that the equilibrium distribution function is proportional to  $e^{-\frac{\epsilon_{\mathbf{p}}}{k_B T}}$ . Choosing the normalization constant to satisfy Eq. (2), we get

$$f^{\rm eq}(\mathbf{p}) = \left(\frac{2\pi}{mk_BT}\right)^{3/2} \frac{N}{V} e^{-\frac{\mathbf{p}^2}{2mk_BT}},\tag{7}$$

which is known as the Maxwell-Boltzmann distribution. Further on we will derive the distribution function (7) within the kinetic theory with no reference to the Gibbs statistical mechanics.

• Substituting the function (7) into Eq. (5), one finds the gas energy density as

$$\varepsilon = \frac{3}{2} k_B T \frac{N}{V}.$$
(8)

- Let us now compute the pressure exerted on the wall of the surface A of the gas container. The pressure results from the collisions of gas particles which are assumed to be elastic. If the particle's momentum before the collision is  $\mathbf{p} = (p_x, p_y, p_z)$  and the surface A is in the yz-plane, the momentum after the collision is  $\mathbf{p}' = (-p_x, p_y, p_z)$ . Therefore, the particle's momentum transferred to the wall is  $2p_x$ .
- One observes that particles with the velocity  $v_x$ , which hit the wall in the time interval dt, are, as shown in Fig. 1, in the volume  $dV = Av_x dt$ . Keeping in mind that the pressure p is a force perpendicular to the surface per unit area and that a force is a momentum transfer in unit time, one finds

$$p = \int_{0}^{\infty} \frac{dp_{x}}{2\pi} \int_{-\infty}^{\infty} \frac{dp_{y}}{2\pi} \int_{-\infty}^{\infty} \frac{dp_{z}}{2\pi} \underbrace{Av_{x}dt}_{dV} \frac{2p_{x}}{Adt} f^{eq}(\mathbf{p})$$
(9)  
$$= 2 \Big(\frac{2\pi}{mk_{B}T}\Big)^{3/2} \frac{N}{V} \int_{0}^{\infty} \frac{dp_{x}}{2\pi} \frac{p_{x}^{2}}{m} e^{-\frac{p_{x}^{2}}{2mk_{B}T}} \Big(\int_{-\infty}^{\infty} \frac{dp_{y}}{2\pi} e^{-\frac{p_{z}^{2}}{2mk_{B}T}}\Big)^{2} = k_{B}T \frac{N}{V},$$

where we have used the integral

$$\int_0^\infty dx \, x^2 e^{-x^2} = \frac{\sqrt{\pi}}{4}.$$
 (10)

The integration over  $p_x$  in the expression (9) is only over positive momenta as we take into account only those particles which are approaching the wall.

# Collisionless kinetic equation

We are going to derive the equation obeyed by the distribution function which is known as <u>kinetic equation</u>, <u>transport equation</u> or <u>Boltzmann equation</u>. Here we will ignore the interparticle collisions but an external potential force will be taken into account.

• At the moment of time t the distribution function is  $f(t, \mathbf{r}, \mathbf{p})$ . After a short interval of time dt the function becomes  $f(t + dt, \mathbf{r} + d\mathbf{r}, \mathbf{p} + d\mathbf{p})$ , which is expanded in the Taylor series as

$$f(t+dt,\mathbf{r}+d\mathbf{r},\mathbf{p}+d\mathbf{p}) = f(t,\mathbf{r},\mathbf{p}) + \frac{\partial f(t,\mathbf{r},\mathbf{p})}{\partial t} dt + \nabla f(t,\mathbf{r},\mathbf{p}) \cdot d\mathbf{r} + \nabla_p f(t,\mathbf{r},\mathbf{p}) \cdot d\mathbf{p}, \quad (11)$$

where  $\nabla_p$  is the momentum gradient.

• Since the inter-particle collisions are ignored but the external potential force **F** is taken into account we have

$$d\mathbf{r} = \mathbf{v}dt = \frac{\mathbf{p}}{m}dt, \qquad d\mathbf{p} = \mathbf{F}dt,$$
 (12)

which substituted into Eq. (11) give

$$f(t+dt,\mathbf{r}+d\mathbf{r},\mathbf{p}+d\mathbf{p}) = f(t,\mathbf{r},\mathbf{p}) + \left(\frac{\partial}{\partial t} + \mathbf{v}\cdot\nabla + \mathbf{F}(\mathbf{r})\cdot\nabla_p\right)f(t,\mathbf{r},\mathbf{p})\,dt.$$
(13)

• As the particles, which at t have been at  $(\mathbf{r}, \mathbf{p})$ , have moved to  $(\mathbf{r} + d\mathbf{r}, \mathbf{p} + d\mathbf{p})$  after dt, the phase space density remains unchanged that is

$$f(t, \mathbf{r}, \mathbf{p}) = f(t + dt, \mathbf{r} + d\mathbf{r}, \mathbf{p} + d\mathbf{p}).$$
(14)

• Using Eq. (11) provides the collisionless kinetic equation or Vlasov equation

$$\left(\frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla + \mathbf{F}(\mathbf{r}) \cdot \nabla_p\right) f(t, \mathbf{r}, \mathbf{p}) = 0.$$
(15)

• If the external force  $\mathbf{F}$  absent, a general solution of the collisionless kinetic equation is

$$f(t, \mathbf{r}, \mathbf{p}) = h(\mathbf{r} - \mathbf{v}t, \mathbf{p}), \tag{16}$$

where  $h(\mathbf{r}, \mathbf{p})$  is an arbitrary function.

• To check that the function (16) solves Eq. (15) one introduce the variable  $\mathbf{s} \equiv \mathbf{r} - \mathbf{v}t$  and computes

$$\left(\frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla\right) f(t, \mathbf{r}, \mathbf{p}) = \left(\frac{\partial s^i}{\partial t} + v^j \frac{\partial s^i}{\partial r^j}\right) \frac{\partial h(\mathbf{s}, \mathbf{p})}{\partial s^i} = \left(-v^i + v^i\right) \frac{\partial h(\mathbf{s}, \mathbf{p})}{\partial s^i} = 0.$$
(17)

#### Gas expansion into vacuum

Let us consider a gas in the thermal equilibrium kept in a container. At a moment of time  $t_0$  the walls of the container are removed and the gas expands into vacuum. How does the process proceed?

• To simplify the problem instead of the container there is a Gaussian distribution of gas density. The initial distribution function is

$$f(t=0,\mathbf{r},\mathbf{p}) = \frac{N}{(\sqrt{2\pi}L_0)^3} \left(\frac{2\pi}{mk_BT}\right)^{3/2} e^{-\frac{\mathbf{r}^2}{2L_0^2}} e^{-\frac{\mathbf{p}^2}{2mk_BT}} = \frac{N}{(mk_BTL_0^2)^{3/2}} e^{-\frac{\mathbf{r}^2}{2L_0^2}} e^{-\frac{\mathbf{p}^2}{2mk_BT}}, \quad (18)$$

where  $L_0$  is the container size.

# Lecture VII

• Knowing the general solution (16), the distribution function for  $t \ge 0$  is written as

$$f(t, \mathbf{r}, \mathbf{p}) = \frac{N}{(mk_B T L_0^2)^{3/2}} e^{-\frac{(\mathbf{r} - \mathbf{v}t)^2}{2L_0^2}} e^{-\frac{\mathbf{p}^2}{2mk_B T}}.$$
(19)

• Taking the position integral of the function (19) we find the momentum distribution

$$\int d^3r f(t, \mathbf{r}, \mathbf{p}) = \left(\frac{2\pi}{mk_B T}\right)^{3/2} e^{-\frac{\mathbf{p}^2}{2mk_B T}},\tag{20}$$

which, as expected, does not change in time.

• Taking the integral over momentum we find the density distribution

$$\rho(t, \mathbf{r}) = \left(\frac{m}{2\pi (mL_0^2 + k_B T t^2)}\right)^{3/2} N \exp\left(-\frac{m \mathbf{r}^2}{2(mL_0^2 + k_B T t^2)}\right),\tag{21}$$

which evolves with time.

- One checks that the density (21) integrated over  $\mathbf{r}$  gives N, as it should.
- Using the formula (21) one shows how the gas expands. The dispersion of the region, where the gas particles are present, is

$$L(t) \equiv \sqrt{\langle \mathbf{r}^2 \rangle - \langle \mathbf{r} \rangle^2} = \sqrt{L_0^2 + \frac{k_B T}{m} t^2} = \sqrt{L_0^2 + \frac{1}{3} v_T^2 t^2},$$
(22)

where we have introduced the quantity  $v_T = \sqrt{\frac{3k_BT}{m}}$  which is the *average velocity* or *thermal velocity* defined through the equality

$$\frac{1}{2}m v_T^2 = \frac{3}{2}k_B T.$$
(23)

As seen, the higher gas temperature the faster expansion.