Gibbs quantum statistical mechanics I

Behavior of atoms and molecules is driven by quantum not classical mechanics. Therefore, a system of many atoms or molecules should be described in terms of quantum mechanics.

Introductory remarks

• A classical description of gases breaks down when a typical de Broglie wavelength of gas constituent λ_B is no longer much smaller than a typical inter-constituent distance d. The latter quantity is related to the gas density as $d = \rho^{-1/3}$ and the de Broglie wavelength estimated for an ideal gas is

$$\lambda_B = \frac{h}{p} = \frac{2\pi\hbar}{\sqrt{2mE}} = \frac{2\pi\hbar}{\sqrt{3mk_BT}}.$$
(1)

Therefore, the gas can be treated as classical if

$$\rho \left(\frac{4\pi^2 \hbar^2}{3mk_B T}\right)^{3/2} \ll 1,\tag{2}$$

otherwise quantum effects are important.

• Measurable quantities – <u>observables</u> – are represented in quantum mechanics by hermitian operators which act in space of states. A goal of quantum statistical mechanics is to compute expectation values of the observables of many-body systems. Having an observable $\hat{\mathcal{O}}$ we are interested in

$$\langle \hat{\mathcal{O}} \rangle \equiv (\psi, \hat{\mathcal{O}}\psi),$$
 (3)

where ψ is the wave function of N particles and (ψ_1, ψ_2) denotes the scalar product of wave functions ψ_1 and ψ_2 .

• Since the system's hamiltonian \hat{H} is assumed to be time independent the wave function ψ can be written as

$$\psi = \sum_{n} C_n \varphi_n. \tag{4}$$

where φ_n are the energy eigenfunctions

$$\hat{H}\varphi_n = E_n\varphi_n,\tag{5}$$

and C_n are complex coefficients which depend on time via $e^{i\frac{E_nt}{\hbar}}$. Using the expansion (4), the expectation value (3) equals

$$\langle \hat{\mathcal{O}} \rangle = \sum_{n,m} C_n^* C_m(\varphi_n, \hat{\mathcal{O}}\varphi_m).$$
(6)

• As we already know, statistical mechanics deals with characteristics averaged over microscopically long interval of time τ

$$\overline{\langle \hat{\mathcal{O}} \rangle} \equiv \frac{1}{\tau} \int_{t}^{t+\tau} dt' \langle \hat{\mathcal{O}} \rangle = \frac{1}{\tau} \sum_{n,m} (\varphi_n, \hat{\mathcal{O}}\varphi_m) \int_{t}^{t+\tau} dt' C_n^* C_m = \sum_{n,m} (\varphi_n, \hat{\mathcal{O}}\varphi_m) \overline{C_n^* C_m}.$$
 (7)

Lecture V

• There is adopted the postulate of random phases

$$\overline{C_n^* C_m} = 0, \quad \text{if} \quad n \neq m.$$
(8)

Thermal environment cancels out an effect of interference of different quantum states.

• Due to the postulate of random phases the expectation value (7) equals

$$\overline{\langle \hat{\mathcal{O}} \rangle} = \sum_{n} \overline{|C_n|^2}(\varphi_n, \hat{\mathcal{O}}\varphi_n).$$
(9)

• Quantum statistical ensemble is defined via the set of φ_n with the coefficients $\overline{|C_n|^2}$.

Microcanonical ensemble

- <u>Microcanonical ensemble</u> is a set of states of isolated system under consideration and all states with the energy E_n from the interval $[U, U + \delta U]$ are equally probable $(|C_n|^2 = 1)$.
- $\Gamma_N(U, V)$ is the number of states φ_n which obey $U \leq E_n \leq U + \delta U$.
- The entropy is defined as in the classical case that is

$$S(U,V) \equiv k_B \ln \Gamma_N(U,V). \tag{10}$$

• At T = 0 the system is in its ground state and $\Gamma_N(U, V)$ equals the ground-state degeneracy \mathcal{N} . Therefore,

$$S(T=0,V) \equiv k_B \ln \mathcal{N}.$$
(11)

• According to the third principle of thermodynamics, the entropy vanishes at T = 0 which means $S(T = 0, V) \ll k_B N$. To fulfill the principle \mathcal{N} cannot be bigger than N^k with $k \ll N$.

Canonical ensemble

• <u>Canonical ensemble</u> is a set of states of a system which is in thermal contact with a heat bath and

$$\overline{|C_n|^2} \sim e^{-\beta E_n},\tag{12}$$

where $\beta \equiv \frac{1}{k_B T}$.

• <u>Partition functions</u> is defined as

$$Q_N(T,V) \equiv \sum_n e^{-\beta E_n}.$$
(13)

• The relation with thermodynamics is established by means of the relations

$$U(T,V) \equiv \frac{1}{Q_N(T,V)} \sum_n E_n e^{-\beta E_n},$$
(14)

$$F(T,V) \equiv -k_B T \ln Q_N(T,V).$$
(15)

where F is the free energy, F = U - TS.

Grand canonical ensemble

• <u>Grand canonical ensemble</u> is a set of states of a system which exchanges heat and particles with its environment and

$$\overline{|C_n|^2} \sim e^{-\beta(E_n - \mu N)},\tag{16}$$

where μ is the chemical potential.

• The grand partition function is

$$\Xi(T, V, \mu) = \sum_{N=0}^{\infty} z^N Q_N(T, V), \qquad (17)$$

where $z \equiv e^{\beta \mu}$.

• The relation with thermodynamics is established by means of the relations

$$U = -\frac{\partial}{\partial\beta} \ln \Xi(T, V, z), \qquad (18)$$

$$\langle N \rangle = z \frac{\partial}{\partial z} \ln \Xi(T, V, z),$$
 (19)

$$pV = k_B T \ln \Xi(T, V, z).$$
⁽²⁰⁾

Einstein's model of crystal

The crystal is treated as a system N independent quantum harmonic oscillators.

- The energy of a single oscillator is $E_n = \hbar \omega \left(n + \frac{1}{2}\right)$, where $n = 0, 1, 2, \ldots$
- The energy of N oscillators is $E_{n_1} + E_{n_2} + \ldots E_{n_N}$ and

$$Q_{N}(T,V) = \sum_{n_{1}=0}^{\infty} \sum_{n_{2}=0}^{\infty} \cdots \sum_{n_{N}=0}^{\infty} e^{-\beta(E_{n_{1}}+E_{n_{2}}+\dots E_{n_{N}})} = \sum_{n_{1}=0}^{\infty} e^{-\beta E_{n_{1}}} \sum_{n_{2}=0}^{\infty} e^{-\beta E_{n_{2}}} \cdots \sum_{n_{N}=0}^{\infty} e^{-\beta E_{n_{N}}} \\ = \left(\sum_{n=0}^{\infty} e^{-\beta\hbar\omega(n+\frac{1}{2})}\right)^{N} = \left(e^{-\frac{\beta\hbar\omega}{2}} \sum_{n=0}^{\infty} e^{-\beta\hbar\omega n}\right)^{N} = \left(\frac{e^{\frac{\beta\hbar\omega}{2}}}{e^{\beta\hbar\omega}-1}\right)^{N},$$
(21)

where the formula

$$\sum_{n=0} q^n = \frac{1}{1-q}$$
(22)

is used.

• The crystal's energy is

$$U = -\frac{\partial}{\partial\beta} \ln Q_N(T, V) = N\left(\frac{\hbar\omega}{e^{\beta\hbar\omega} - 1} + \frac{\hbar\omega}{2}\right),\tag{23}$$

where the second term corresponds to the so-called *zero-point fluctuations*.

Lecture V

• The heat capacity equals

$$C_V = \left(\frac{\partial U}{\partial T}\right)_V = \frac{N}{k_B T^2} \frac{(\hbar\omega)^2 e^{\beta\hbar\omega}}{(e^{\beta\hbar\omega} - 1)^2},\tag{24}$$

which can be approximated as

$$C_V \approx \begin{cases} Nk_B & \text{for} \quad k_B T \gg \hbar \omega, \\ \frac{N(\hbar \omega)^2}{k_B T^2} e^{-\beta \hbar \omega} & \text{for} \quad k_B T \ll \hbar \omega. \end{cases}$$
(25)

The heat capacity as a function of temperature is shown in Fig. 1.

- When $T \to 0$ the heat capacity vanishes in agreement with the third principle of thermodynamics.
- If the oscillations occur in three dimensions, N should be replaced by 3N.

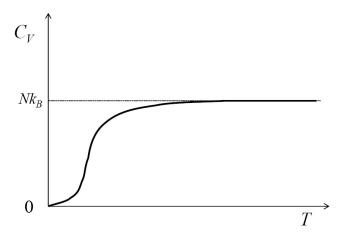


Figure 1: Heat capacity as a function of temperature