

Letter

Correspondence of multiplicity and energy distributions

Maciej Rybczyński^{1,a}, Zbigniew Włodarczyk¹

¹ Institute of Physics, Jan Kochanowski University, 25-406 Kielce, Poland

Received: 20 October 2020 / Accepted: 20 November 2020 © Società Italiana di Fisica and Springer-Verlag GmbH Germany, part of Springer Nature 2021 Communicated by Ulf Meissner

Abstract The evaluation of the number of ways we can distribute energy among a collection of particles in a system is important in many branches of modern science. In particular, in multiparticle production processes the measurements of particle yields and kinematic distributions are essential for characterizing their global properties and to develop an understanding of the mechanism for particle production. We demonstrate that energy distributions are connected with multiplicity distributions by their generating functions.

For the count probability distribution, P(N), the generating function G(z) is defined as:

$$G(z) = \sum_{N=0}^{\infty} P(N) z^{N}.$$
 (1)

Thus far the dummy variable z of the generating function has been considered just as a technical auxiliary variable ("book keeping variable"). Only in the so called method of *collective marks* one gives a probability interpretation for the variable z^{-1} . If we mark each of the N elements in the set independently with probability 1-z and leave it unmarked with probability z, then G(z) is the probability that there is no mark in the whole set.

In this letter multiplicity distributions P(N) in quasi power-law ensembles and their generating functions G(z) are discussed. They are connected with the energy distributions F(E) of elements in the ensemble.

Note, that generating functions of NBD and BD (shown in Table 1) are in fact some quasi-power functions of z and as such can be written in the form of the corresponding Tsallis distributions [5–8].

Published online: 02 January 2021

$$G(z) = \exp_q \left[\langle N \rangle (1 - z) \right]$$

$$= \left[1 + (q - 1) \langle N \rangle (1 - z) \right]^{\frac{1}{1 - q}}, \qquad (2)$$

where q - 1 = 1/k for NBD, q - 1 = -1/K for BD, and $q - 1 \rightarrow 0$ for PD. For

$$z = 1 - \frac{E}{U} \tag{3}$$

with the total available energy

$$U = \sum_{i=1}^{N} E_i,\tag{4}$$

the multiplicity generating function (2) gives the energy distribution

$$F(E) = G(z = 1 - E/U) = \left[1 + (q - 1)\frac{E}{T}\right]^{\frac{1}{1 - q}} \tag{5}$$

which is the well known Tsallis distribution [5], and which for $q \to 1$ becomes Boltzmann-Gibbs distribution. This distribution was first proposed in [9,10] as the simplest formula extrapolating exponential behavior observed for low transverse momenta to power law behavior at large transverse momenta. At present it is known as the QCD-inspired *Hagedorn formula* [11,12]. Function (5) is usually interpreted in terms of the statistical model of particle production employing the Tsallis non-extensive statistics [5–7] and widely used in description of multiparticle production processes [13,14] 2 .

To explain the correspondence of multiplicity and energy distributions (schematically illustrated in Fig. 1), let us consider a simple example. For fixed number of particles N, energy distribution emerges directly from the calculus of probability for a situation known as *induced partition* [15]. In short: N-1 randomly chosen independent points

² For an updated bibliography on this subject, see http://tsallis.cat.cbpf.br/biblio.htm.



¹ The method of collective marks was originated by van Dantzig [1], and discussed in [2] and [3]. Recently, the collective marks method was used to find the probability generating function for first passage probabilities of Markov chains [4].

^a e-mail: maciej.rybczynski@ujk.edu.pl (corresponding author)

3 Page 2 of 4 Eur. Phys. J. A (2021) 57:3

Table 1 Distributions P(N) used in this work: Poisson (PD), Negative Binomial (NBD) and Binomial (BD) and their generating functions G(z)

	P (N)	$G\left(z\right)$
PD	$\frac{\lambda^N}{N!} \exp\left(-\lambda\right)$	$\exp\left[\lambda\left(z-1\right)\right]$
NBD	$\frac{\Gamma(N+k)}{\Gamma(N+1)\Gamma(k)}p^N(1-p)^k$	$\left[1 - \frac{p}{1-p}\left(z - 1\right)\right]^{-k}$
BD	$\frac{K!}{N!(K-N)!} p^N (1-p)^{K-N}$	$[1+p(z-1)]^K$

 $\{U_1, \ldots, U_{N-1}\}$ split a segment (0, U) into N parts, whose length is distributed according to:

$$F(E|N) = \frac{N-1}{U} \left(1 - \frac{E}{U} \right)^{N-2}.$$
 (6)

The length of the kth part corresponds to the value of energy $E_k = U_{k+1} - U_k$ (for ordered U_k). Whereas for fixed N one have (6), then for N fluctuating according to P(N), the resulting energy distribution is

$$F(E) = \sum_{N=2}^{\infty} P(N) F(E|N).$$
 (7)

For P(N) given by BD, PD, and NBD, Eq. (7) leads to Tsallis distribution given by equation (5). Relationships between Poissonian multiplicity distribution and Boltzmann-Gibbs energy distribution are discussed in more detail in the Appendix.

Note that P(N), defined for N > 1, describe multiplicity distribution in the full phase-space. In experiments, particle multiplicity is measured usually only within some window of phase-space. Let us assume that the detection process is a Bernoulli process described by the BD (K = 1 and $p = \alpha$ for a fixed experimental acceptance $\alpha < 1$). The number of registered particles is

$$M = \sum_{i=1}^{N} n_i,\tag{8}$$

where n_i follows the BD with the generating function $G_{BD}(z)$ and N comes from P(N) with the generating function G(z). The measured multiplicity distribution

$$P(M) = \frac{1}{M!} \frac{d^M H(z)}{dz^M} \bigg|_{z=0}$$
 (9)

is therefore given by generating function $H(z) = G(G_{BD}(z))$. Such rough procedure applied to NBD, BD or PD gives again the same distributions but with modified parameters: $p \to \alpha p/[1-p(1-\alpha)]$ for NBD, $p \to \alpha p$ for BD, and $\lambda \to \alpha \lambda$ for PD. The measured multiplicity

distribution is given by

$$P(M) = \sum_{N=M}^{\infty} P(N) P(M|N)$$
(10)

with the acceptance function

$$P(M|N) = \frac{N!}{M!(N-M)!} \alpha^{M} (1-\alpha)^{N-M}$$
 (11)

Detection process extend P(M) distribution to multiplicities M = 0 and M = 1, namely: $P(0) = \sum_{N=2}^{\infty} P(N) (1 - \alpha)^N$ and $P(1) = \sum_{N=2}^{\infty} P(N) N\alpha (1 - \alpha)^{N-1}$.

The statistical properties of the energy division between a set of particles are completely characterized by the generating function G(z). Despite correspondence between multiplicity and energy distributions, the multiplicity distribution gives in practice complementary information to the energy distribution, because P(N) is defined by the N^{th} derivative of G(z) = F(E) at E = U, i.e., in the region not available experimentally in measurements at collider experiments 3 .

The above considerations (in particular equality given by equation (5) apply to a single statistical ensembles (as realized in hadronic collisions). In nuclear collisions there are usually many statistical systems, independent from one another. In superposition models of hadron production, the number of particles N, as registered in the experiment, is composed from independent production from N_S sources [16]. For a fixed number of sources (neglecting the nuclear modification factor) we have $F_{AA}(E) = N_S \cdot F_{pp}(E)$ and $G_{AA}(z) = (G_{pp}(z))^{N_S}$, what results in equality:

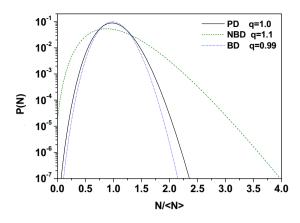
$$F_{AA}(E) = N_S \cdot (G_{AA}(z = 1 - E/U_{pp}))^{1/N_S}.$$
 (12)

For fluctuating numbers of sources, the resulting multiplicity distribution is given by the compound distribution defined by generating function $G_{AA}(z) = H(G_{pp}(z))$, where H(z) is the generating function of distribution of the number of sources. In this case we have a relationship $F_{AA}(E) = \langle N_S \rangle H^{-1}[G_{AA}(z)]$, where H^{-1} is the inverse function to H(z), what is troublesome in practical applications.



³ Similarly as N^{th} derivatives of G(z) taken at z=0 define multiplicity distribution P(N), the respective derivatives taken at z=1 define factorial moments \mathcal{F}_N . Derivatives of $\ln(G(z))$ taken at z=0 and z=1 define combinants \mathcal{C}_N and cumulant factorial moments \mathcal{K}_N , respectively.

Eur. Phys. J. A (2021) 57:3 Page 3 of 4 3



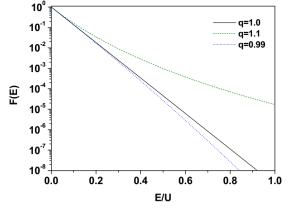


Fig. 1 Multiplicity distributions and corresponding energy distributions

This research was supported by the Polish National Science Centre grant 2016/23/B/ST2/00692 (MR).

Data Availability Statement This manuscript has no associated data or the data will not be deposited. [Authors' comment: The work presented here is theoretical and all the required formulas are given in the article. All data generated or analysed during this study are included in this published article.]

Appendix A: Boltzmann-Gibbs energy distribution and Poissonian multiplicity distribution

Suppose that one has N independently produced particles with energies $\{E_{1,\dots,N}\}$, distributed according to Boltzmann distribution,

$$F(E) = \frac{1}{T} \exp\left(-\frac{E}{T}\right) \tag{A.1}$$

with "temperature" parameter $T=\langle E\rangle$. The sum of energies, $U=\sum_{i=1}^N E_i$ is then distributed according to gamma distribution

$$F_N(U) = \frac{1}{T(N-1)!} \left(\frac{U}{T}\right)^{N-1} \exp\left(-\frac{U}{T}\right)$$
$$= F_{N-1}(U) \frac{U}{N-1}$$
(A.2)

with cumulative distribution equal to:

$$F_N(>U) = 1 - \sum_{i=1}^{N-1} \frac{1}{(i-1)!} \left(\frac{U}{T}\right)^{i-1} \exp\left(-\frac{U}{T}\right).$$
 (A.3)

Looking for such N that $\sum_{i=0}^{N} E_i \le U \le \sum_{i=0}^{N+1} E_i$ we find its distribution. which has known Poissonian form

$$P(N) = F_{N+1}(> U) - F_N(> U)$$

$$= \frac{(U/T)^N}{N!} \exp\left(-\frac{U}{T}\right)$$

$$= \frac{\langle N \rangle^N}{N!} \exp\left(-\langle N \rangle\right) \tag{A.4}$$

with $\langle N \rangle = U/T$.

For the constrained systems (if the available energy is limited, U = const), whenever we have independent variables $\{E_{1,\dots,N}\}$ taken from the exponential distribution (A.1), the corresponding multiplicity N has Poissonian distribution (A.4)⁴. However, if the multiplicity is limited, N = const, the resulting *conditional probability* becomes:

$$F(E|N) = \frac{F_1(E) F_{N-1}(U - E)}{F_N(U)}$$
$$= \frac{N-1}{U} \left(1 - \frac{E}{U}\right)^{N-2}$$
(A.5)

the same as given by Eq. (5), and only in the limit $N \to \infty$ the energy distribution goes to the Boltzmann distribution (A.1). For fluctuating multiplicity according to Poisson distribution, the energy distribution is given by (A.1).

In the same way, as demonstrated in Ref. [17], Tsallis energy distribution is connected with the NBD of multiplicity.

References

- D. Van Dantzig, Colloques internationaux du CNRS 13, 29–45 (1949)
- J.T. Runnenburg, On the use of collective marks in queueing theory, in *Congestion theory*, ed. by W.L. Smith, W.E. Wilkinson (University of North Carolina Press, Chapel Hill, 1965), pp. 399–438
- L. Kleinrock, Queueing systems, Chapter 7, vol. 1 (Wiley, New York, 1975)

⁴ Actually this is the method of generating Poisson distribution in the numerical Monte Carlo codes.



3 Page 4 of 4 Eur. Phys. J. A (2021) 57:3

- 4. Y. Zhang, M. Hlynka, P.H. Brill, arXiv:1908.04370v1 [math.PR]
- C. Tsallis, J. Statist. Phys. 52, 479 (1988). https://doi.org/10.1007/ BF01016429
- C. Tsallis, Eur. Phys. J. A 40, 257 (2009a). https://doi.org/10.1140/epja/i2009-10799-0. [arXiv:0812.4370 [physics.data-an]]
- 7. C. Tsallis, Introduction to nonextensive statistical mechanics (Springer, Berlin, 2009)
- G. Wilk, Z. Włodarczyk, Acta Phys. Polon. B 46(6), 1103 (2015). https://doi.org/10.5506/APhysPolB.46.1103. [arXiv:1501.01936 [cond-mat.stat-mech]]
- C. Michael, L. Vanryckeghem, J. Phys. G 3, L151 (1977). https://doi.org/10.1088/0305-4616/3/8/002
- C. Michael, Prog. Part. Nucl. Phys. 2, 1 (1979). https://doi.org/10. 1016/0146-6410(79)90002-4
- G. Arnison et al., UA1 collaboration. Phys. Lett. 118B, 167 (1982). https://doi.org/10.1016/0370-2693(82)90623-2

- R. Hagedorn, Riv. Nuovo Cim. 6N10, 1 (1983). https://doi.org/10. 1007/BF02740917
- C.Y. Wong, G. Wilk, L.J.L. Cirto, C. Tsallis, Phys. Rev. D 91(11), 114027 (2015). https://doi.org/10.1103/PhysRevD.91. 114027. [arXiv:1505.02022 [hep-ph]]
- G. Wilk, Z. Wlodarczyk, Eur. Phys. J. A 48, 161 (2012). https://doi.org/10.1140/epja/i2012-12161-y. [arXiv:1203.4452 [hep-ph]]
- 15. W. Feller, An introduction to probability theory and its applications, vol. II (Wiley, New York, 1966)
- W. Broniowski, A. Olszewski, Phys. Rev. C 95(6), 064910 (2017). https://doi.org/10.1103/PhysRevC.95.064910. [arXiv:1704.01532 [nucl-th]]
- G. Wilk, Z. Wlodarczyk, Phys. A 376, 279 (2007). https://doi.org/ 10.1016/j.physa.2006.10.042

