

Boson gas

In this lecture the method of thermal field theory will be applied to a weakly interacting boson gas.

Partition function

- We consider a boson gas described the Lagrangian density

$$\mathcal{L}(x) = \frac{1}{2} \partial^\mu \phi(x) \partial_\mu \phi(x) - \frac{1}{2} m^2 \phi^2(x) - \frac{\lambda}{4!} \phi^4(x), \quad (1)$$

where $\phi(x)$ is the real scalar, m is the mass parameter and λ is the coupling constant.

- As we remember from the Lecture II, the partition function defined as

$$Z \equiv \sum_{\alpha} \langle \alpha | e^{-\beta \hat{H}} | \alpha \rangle, \quad (2)$$

where $\beta \equiv T^{-1}$ is the inverse temperature and \hat{H} is the system's Hamiltonian.

- The partition function of non-interacting boson gas derived in Lecture II is

$$Z_0 = \exp \left[-V \int \frac{d^3 k}{(2\pi)^3} \ln(1 - e^{-\beta \omega_{\mathbf{k}}}) \right], \quad (3)$$

where V is the system's volume and $\omega_{\mathbf{k}} \equiv \sqrt{\mathbf{p}^2 + m^2}$.

- One computes the partition function as

$$\ln Z_0 = -\frac{V}{2\pi^2} \int_0^\infty dk k^2 \ln(1 - e^{-\beta \sqrt{m^2 + k^2}}) = -\frac{V}{6\pi^2} \int_0^\infty dk \frac{dk^3}{dk} \ln(1 - e^{-\beta \sqrt{m^2 + k^2}}). \quad (4)$$

- Performing the partial integration, one finds

$$\ln Z_0 = \frac{V}{6\pi^2 T} \int_0^\infty \frac{dk k^4}{\sqrt{m^2 + k^2}} \frac{1}{e^{\beta \sqrt{m^2 + k^2}} - 1}. \quad (5)$$

- Further on, we will be mostly interested in the hot gas such that $T \gg m$. Then, the bosons can be treated as massless and for $m = 0$ we will deal with simple analytical formulas. The partition function equals

$$\ln Z_0 = \frac{V T^3}{6\pi^2} \underbrace{\int_0^\infty \frac{dx x^3}{e^x - 1}}_{=\frac{\pi^4}{15}} = \frac{\pi^2 V T^3}{90}. \quad (6)$$

- The system's energy U , free energy $F = U - TS$ and pressure p are

$$U \equiv -\frac{d}{d\beta} \ln Z(T), \quad F \equiv -T \ln Z(T), \quad p = -\left(\frac{\partial F}{\partial V} \right)_T. \quad (7)$$

- The partition function (6) gives

$$U_0 = \frac{\pi^2 V T^4}{30}, \quad F_0 = -\frac{\pi^2 V T^4}{90}, \quad p_0 = \frac{\pi^2 T^4}{90}. \quad (8)$$

First order correction to partition function

- As we remember from the Lecture III, the partition function, which is of the form appropriate for perturbative expansion, is

$$Z = \text{Tr} [e^{-\beta \hat{H}^0} \mathcal{T}[e^{-\int_0^\beta d\tau \hat{H}_{\text{int}}^I(-i\tau)}]]. \quad (9)$$

- The zeroth order contribution to $\mathcal{T}[e^{-\int_0^\beta d\tau \hat{H}_{\text{int}}^I(-i\tau)}]$ is unity and the first order contribution, which corresponds to the diagram shown in Fig. 1, equals

$$\begin{aligned} Z_{(1)} &= \text{Tr} [e^{-\beta \hat{H}^0} \mathcal{T}[e^{-\int_0^\beta d\tau \hat{H}_{\text{int}}^I(-i\tau)}]]_{(1)} = \text{Tr} [e^{-\beta \hat{H}^0}] \frac{\text{Tr} [e^{-\beta \hat{H}^0} \mathcal{T}[e^{-\int_0^\beta d\tau \hat{H}_{\text{int}}^I(-i\tau)}]]}{\text{Tr} [e^{-\beta \hat{H}^0}]} \Big|_{(1)} \\ &= -Z_0 \frac{\lambda}{8} \int_0^\beta d^4x \Delta(0) \Delta(0) = -Z_0 \frac{\lambda}{8} (\Delta(0))^2 \int_0^\beta d^4x. \end{aligned} \quad (10)$$

- As we remember, the function $\Delta(0)$ is identified with the function $\Delta^>(0)$. Since the function $\Delta^>(\tau, \mathbf{x})$ is, see Eq. (24) of Lecture III,

$$\Delta^>(\tau, \mathbf{x}) = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{k}}} \left[(f(\omega_{\mathbf{k}}) + 1) e^{-\omega_{\mathbf{k}}\tau} e^{i\mathbf{k}\cdot\mathbf{x}} + f(\omega_{\mathbf{k}}) e^{\omega_{\mathbf{k}}\tau} e^{-i\mathbf{k}\cdot\mathbf{x}} \right], \quad (11)$$

where the boson distribution function equals

$$f(\omega_{\mathbf{k}}) \equiv \frac{1}{e^{\beta\omega_{\mathbf{k}}} - 1}, \quad (12)$$

the function $\Delta^>(0)$ is

$$\Delta^>(0) = \int \frac{d^3k}{(2\pi)^3} \frac{2f(\omega_{\mathbf{k}}) + 1}{2\omega_{\mathbf{k}}}. \quad (13)$$

- After performing the trivial angular integral, Eq. (13) becomes

$$\Delta^>(0) = \frac{1}{4\pi^2} \int_0^\infty \frac{dk k^2}{\sqrt{m^2 + k^2}} \frac{1 + e^{-\beta\sqrt{m^2+k^2}}}{1 - e^{-\beta\sqrt{m^2+k^2}}}. \quad (14)$$

- Since for $k \gg m$ and $k \gg T$ the integrand linearly grows with k , the integral in Eq. (14) is quadratically divergent. One observes that the divergence remains in the zero temperature limit that is when $\beta \rightarrow \infty$. Therefore, it is the ultraviolet divergence which is well known in vacuum QFT.
- To get a finite result one should subtract the vacuum contribution from the formula (13). Since $f(\omega_{\mathbf{k}}) = 0$ in vacuum, the subtraction is done as follows

$$\Delta_R^>(0) = \int \frac{d^3k}{(2\pi)^3} \frac{2f(\omega_{\mathbf{k}}) + 1}{2\omega_{\mathbf{k}}} - \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_{\mathbf{k}}} = \int \frac{d^3k}{(2\pi)^3} \frac{f(\omega_{\mathbf{k}})}{\omega_{\mathbf{k}}} \quad (15)$$

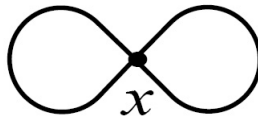


Figure 1: The first order contribution to the partition function (9)

and the renormalized Green's function, which is finite, equals

$$\Delta_R^>(0) = \frac{1}{2\pi^2} \int_0^\infty \frac{dk k^2}{\sqrt{m^2 + k^2}} \frac{1}{e^{\beta\sqrt{m^2+k^2}} - 1}. \quad (16)$$

- Assuming that $m = 0$, one finds

$$\Delta_R^>(0) = \frac{1}{2\pi^2} \int_0^\infty \frac{dk k}{e^{\beta k} - 1} = \frac{T^2}{2\pi^2} \underbrace{\int_0^\infty \frac{dx x}{e^x - 1}}_{=\frac{\pi^2}{6}} = \frac{T^2}{12}. \quad (17)$$

- The four-dimensional integral over x is

$$\int_0^\beta d^4x \equiv \int_0^\beta d\tau \int d^3x = \beta V, \quad (18)$$

where V is the system's volume.

- The first order correction to the partition function is

$$Z_{(1)} = -Z_0 \frac{\lambda}{1152} V T^3. \quad (19)$$

- Using the expression (6) of Z_0 , the partition function is found as

$$Z = \exp\left(\frac{\pi^2 V T^3}{90}\right) \left[1 - \frac{\lambda}{1152} V T^3\right]. \quad (20)$$

- Since the second term in the square bracket in Eq. (20) should be, as a perturbative correction, much smaller than unity, the expression in the bracket can be approximated as

$$1 - \frac{\lambda}{1152} V T^3 \approx \exp\left(-\frac{\lambda}{1152} V T^3\right), \quad (21)$$

which allows one to rewrite the partition function (20) in the following form

$$\ln Z = \frac{\pi^2 V T^3}{90} - \frac{\lambda}{1152} V T^3 = \frac{\pi^2 V T^3}{90} \left[1 - \frac{5\lambda}{64\pi^2}\right]. \quad (22)$$

- The energy, free energy and pressure, which include the first order corrections, are

$$U = \frac{\pi^2 V T^4}{30} \left[1 - \frac{5\lambda}{64\pi^2}\right], \quad (23)$$

$$F = -\frac{\pi^2 V T^4}{90} \left[1 - \frac{5\lambda}{64\pi^2}\right], \quad (24)$$

$$p = \frac{\pi^2 T^4}{90} \left[1 - \frac{5\lambda}{64\pi^2}\right]. \quad (25)$$

- Using the apparatus of thermal field theory, we have managed to go beyond the ideal gas approximation.
- Needless to say, the procedure of perturbative expansion can be systematically extended to higher orders.