

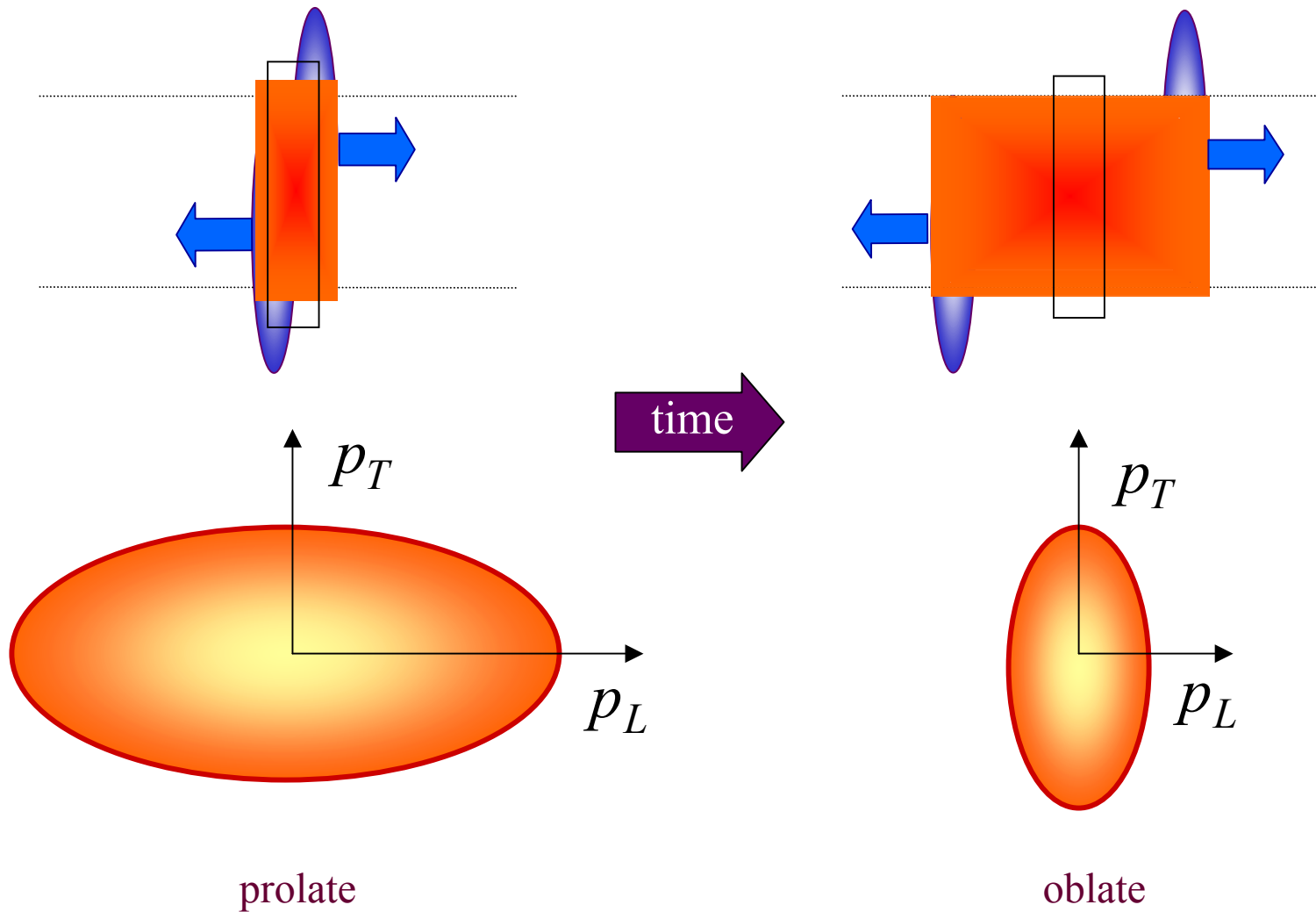
# Parton Momentum Distribution Prior to Equilibrium

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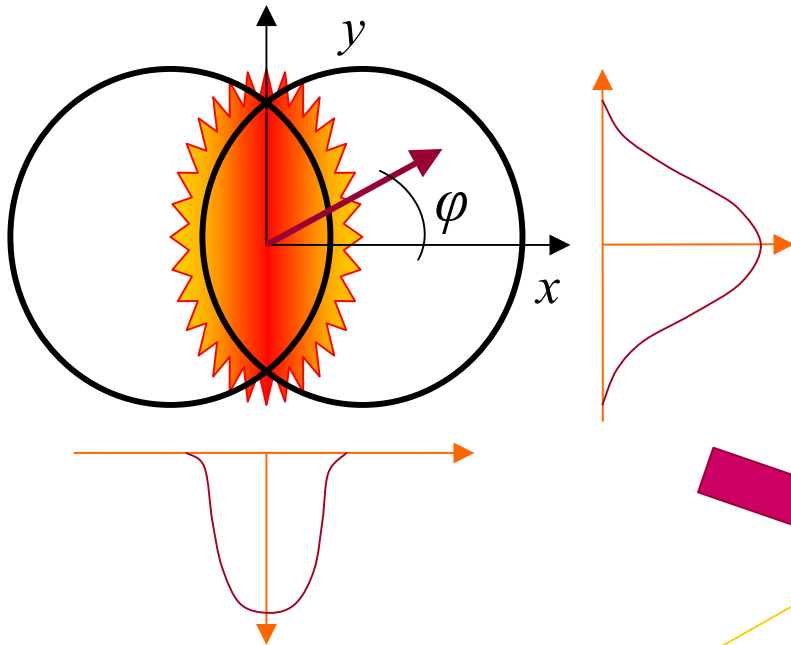
*in collaboration with Weronika Jas*

# Evolution of Parton Momentum Distribution



# Elliptic Flow & Equilibration

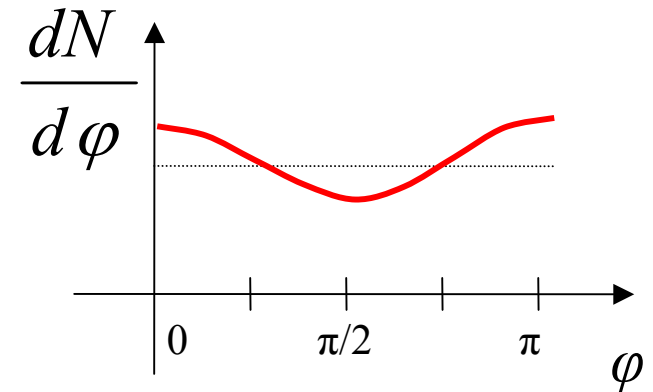
Success of hydrodynamic models in describing elliptic flow



Hydrodynamics

$$\left( \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla \right) \mathbf{v} = - \frac{\nabla p}{\rho}$$

Hydrodynamics requires local thermodynamical equilibrium!



# Equilibration Time

$$v_2 \sim \varepsilon = \frac{\langle y^2 \rangle - \langle x^2 \rangle}{\langle y^2 \rangle + \langle x^2 \rangle}$$

**Eccentricity decays due to the free streaming!**

$$\varepsilon \searrow \Rightarrow v_2 \searrow$$



$$t_{\text{eq}} \leq 1 \text{ fm}/c$$

time of equilibration

# Decay of Eccentricity

free-streaming model

$$f(\mathbf{p}, \mathbf{r}, t) \sim \exp \left[ -\frac{(x - v_x t)^2}{2\sigma_x^2} - \frac{(y - v_y t)^2}{2\sigma_y^2} - \frac{(z - v_z t)^2}{2\sigma_z^2} - \frac{Y^2}{2\Delta Y^2} \right] \frac{P(p_T)}{p_T^2 \text{ch} Y}$$

$$\left\{ \begin{aligned} \langle x^2 \rangle &= \int \frac{d^3 p}{(2\pi)^3} \int d^3 r x^2 f(\mathbf{p}, \mathbf{r}, t) = \sigma_x^2 + \alpha t^2 \\ \langle y^2 \rangle &= \int \frac{d^3 p}{(2\pi)^3} \int d^3 r y^2 f(\mathbf{p}, \mathbf{r}, t) = \sigma_y^2 + \alpha t^2 \end{aligned} \right.$$

$$m=0$$

$$\varepsilon(t) = \frac{\langle y^2 \rangle - \langle x^2 \rangle}{\langle y^2 \rangle + \langle x^2 \rangle} = \frac{\varepsilon(0)}{1 + \frac{\alpha}{R_T^2} t^2}$$

$$\varepsilon(0) = \frac{\sigma_y^2 - \sigma_x^2}{\sigma_y^2 + \sigma_x^2}, \quad R_T^2 = \frac{\sigma_y^2 + \sigma_x^2}{2}$$

$$\alpha = 1/2 \text{ at } Y \approx 0$$

## Decay of Eccentricity cont.

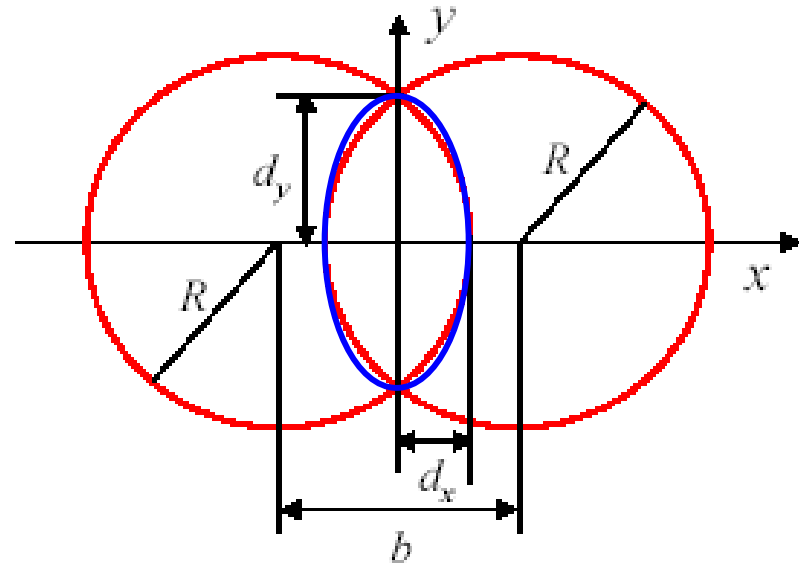
$$\frac{x^2}{d_x^2} + \frac{y^2}{d_y^2} \leq 1$$

$$d_x^2 = (R - b/2)^2$$

$$d_y^2 = R^2 - b^2 / 4$$

$$\sigma_x = \sqrt{\langle x^2 \rangle} = d_x / 2$$

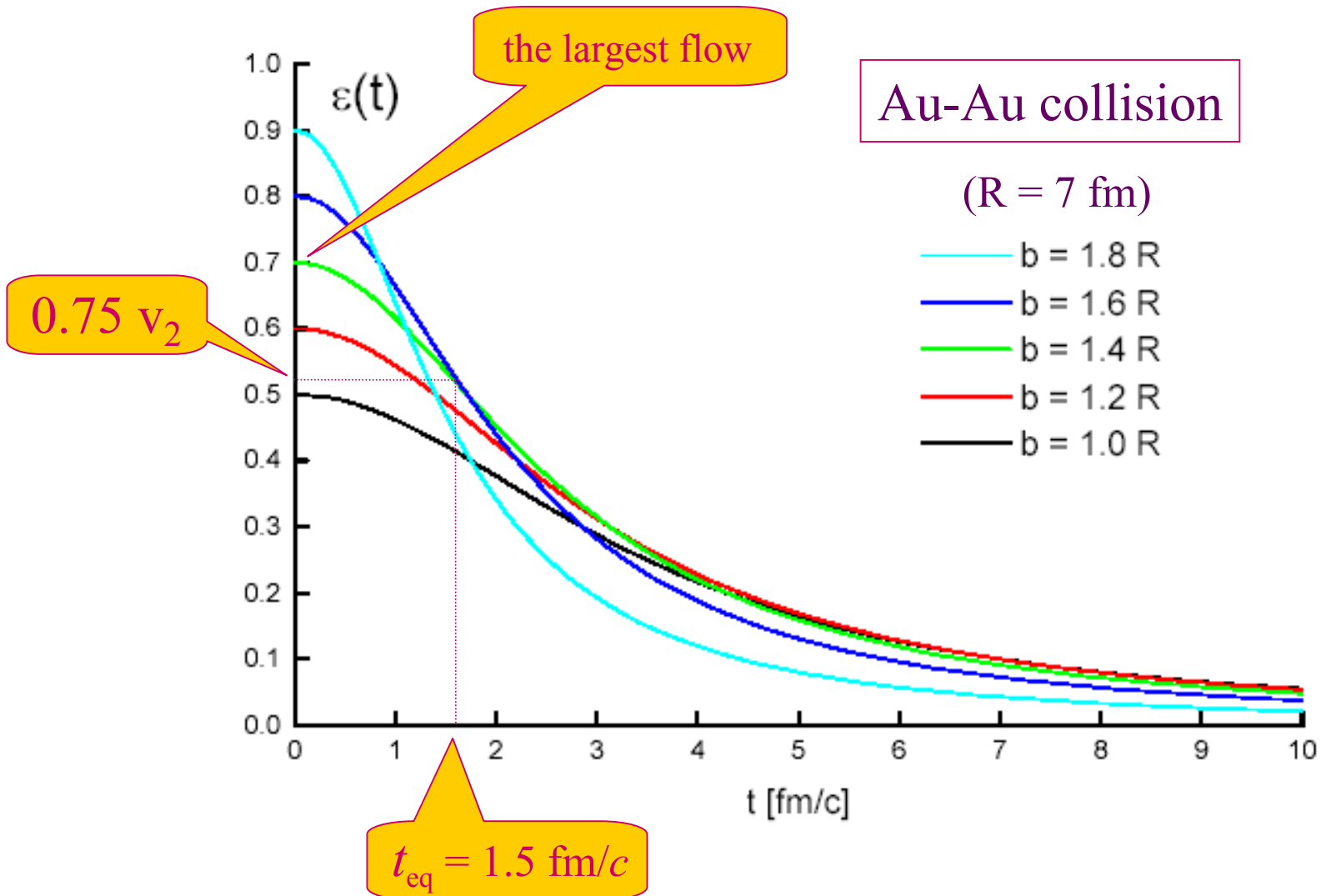
$$\sigma_y = \sqrt{\langle y^2 \rangle} = d_y / 2$$



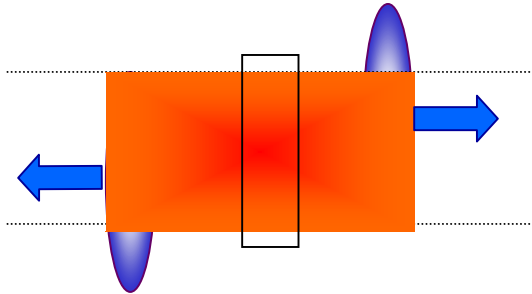
$$\varepsilon(0) = \frac{b}{2R}$$

$$R_T^2 = \frac{R^2}{4} (1 - \varepsilon(0))$$

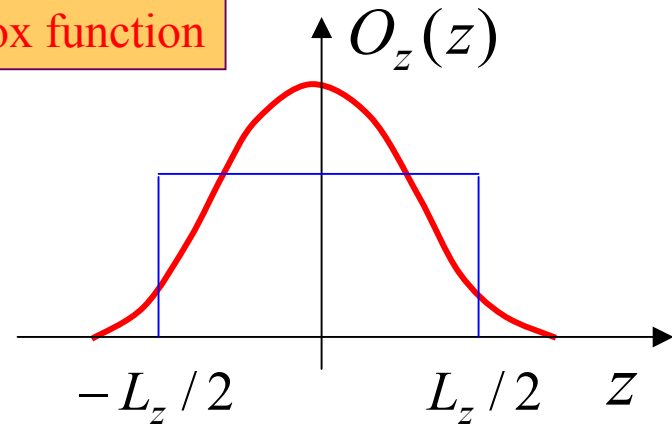
# Decay of Eccentricity cont.



# Momentum Distribution in a Box



box function



$$\int_{-L_z/2}^{L_z/2} dz = \int_{-\infty}^{+\infty} dz O_z(z), \quad \int_{-L_z/2}^{L_z/2} dz z^2 = \int_{-\infty}^{+\infty} dz z^2 O_z(z)$$

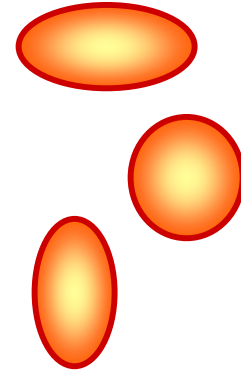
$$O(\mathbf{r}) = O_x(x) O_y(y) O_z(z) = \left(\frac{6}{\pi}\right)^{3/2} \exp\left[-\frac{6x^2}{L_x^2} - \frac{6y^2}{L_y^2} - \frac{6z^2}{L_z^2}\right]$$



# Evolution of Momentum Distribution

$$\rho(t) \equiv \frac{2\langle p_L^2 \rangle}{\langle p_T^2 \rangle}$$

$$\left\{ \begin{array}{ll} \rho > 1 & \text{prolate} \\ \rho = 1 & \text{spherical} \\ \rho < 1 & \text{oblate} \end{array} \right.$$

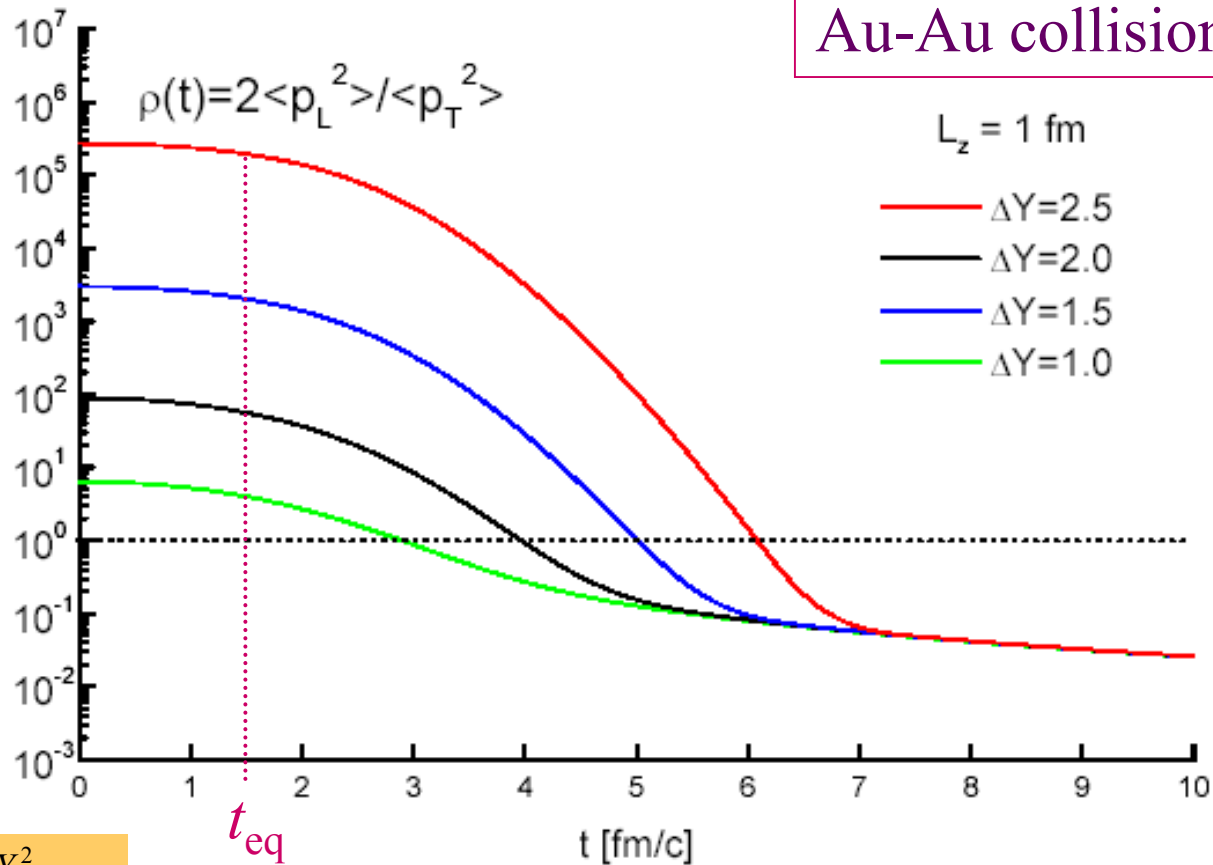


$$\left\{ \begin{array}{l} \langle p_i^2 \rangle = \frac{1}{N} \int d^3 r O(\mathbf{r}) \int \frac{d^3 p}{(2\pi)^3} p_i^2 f(\mathbf{p}, \mathbf{r}, t), \quad i = L, T \\ N = \int d^3 r O(\mathbf{r}) \int \frac{d^3 p}{(2\pi)^3} f(\mathbf{p}, \mathbf{r}, t) \end{array} \right.$$

# Evolution of Momentum Distribution cont.

$$\sigma_x = \sigma_y = L_x = L_y = 3 \text{ fm}, \quad \sigma_z = L_z = 1 \text{ fm}$$

Au-Au collision

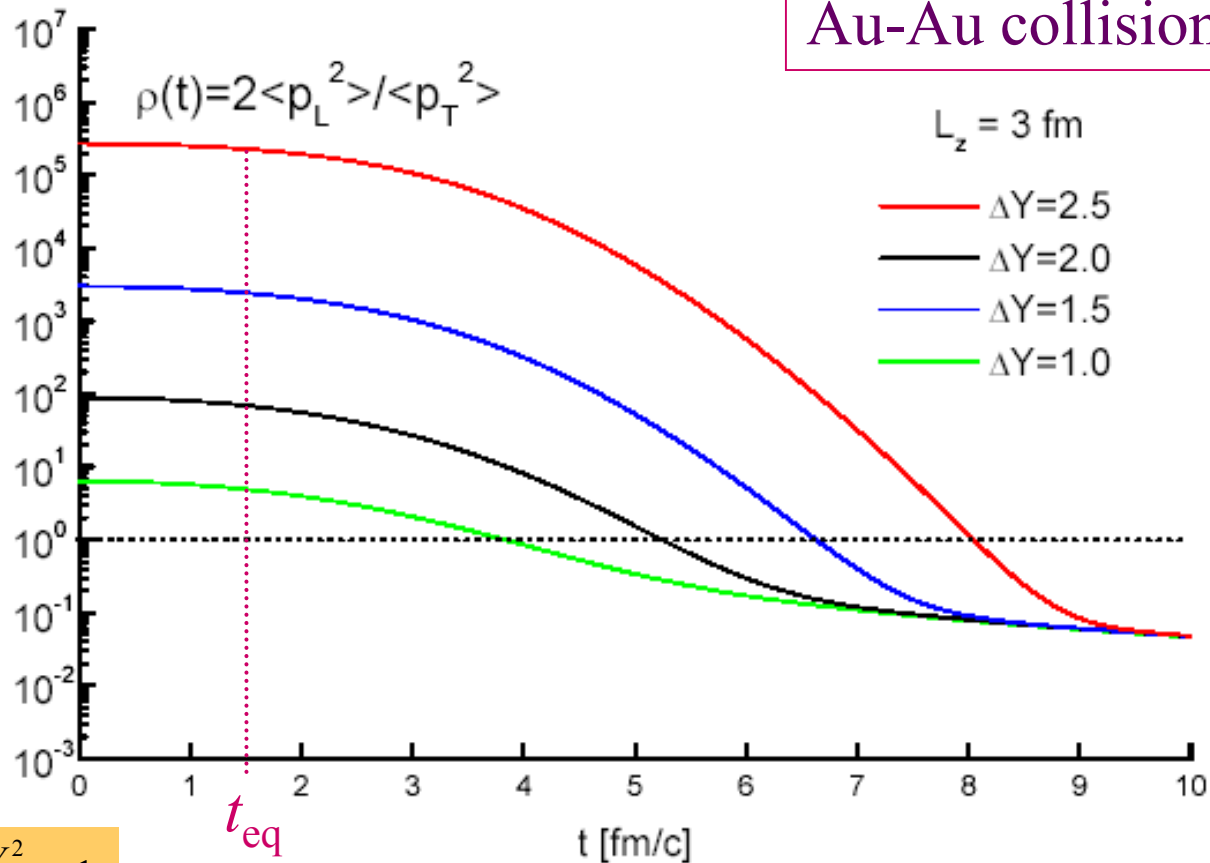


$$\rho(0) = e^{2\Delta Y^2} - 1$$

# Evolution of Momentum Distribution cont.

$$\sigma_x = \sigma_y = L_x = L_y = 3 \text{ fm}, \quad \sigma_z = 1 \text{ fm}, \quad L_z = 3 \text{ fm},$$

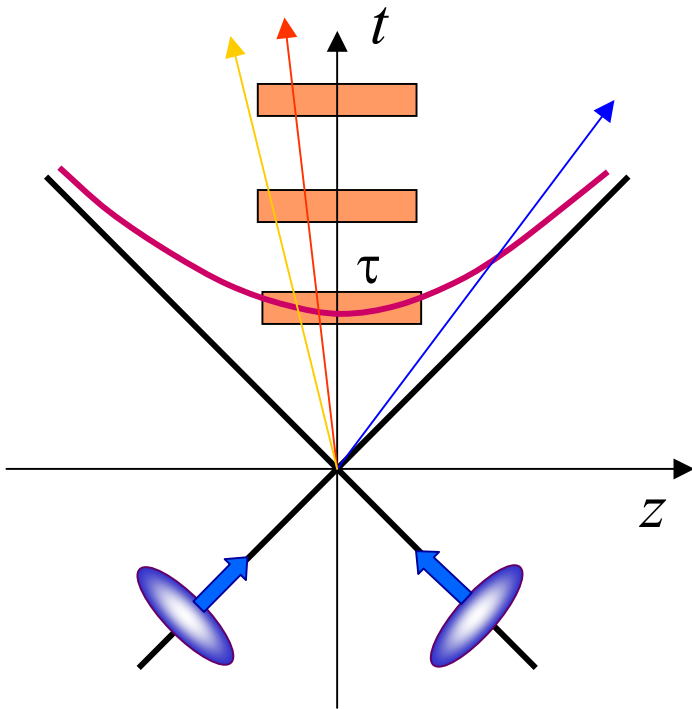
Au-Au collision



$$\rho(0) = e^{2\Delta Y^2} - 1$$

# Effect of finite formation time

$$f(\mathbf{p}, \mathbf{r}, t) \rightarrow \Theta(t\sqrt{1 - v_z^2} - \tau) f(\mathbf{p}, \mathbf{r}, t)$$

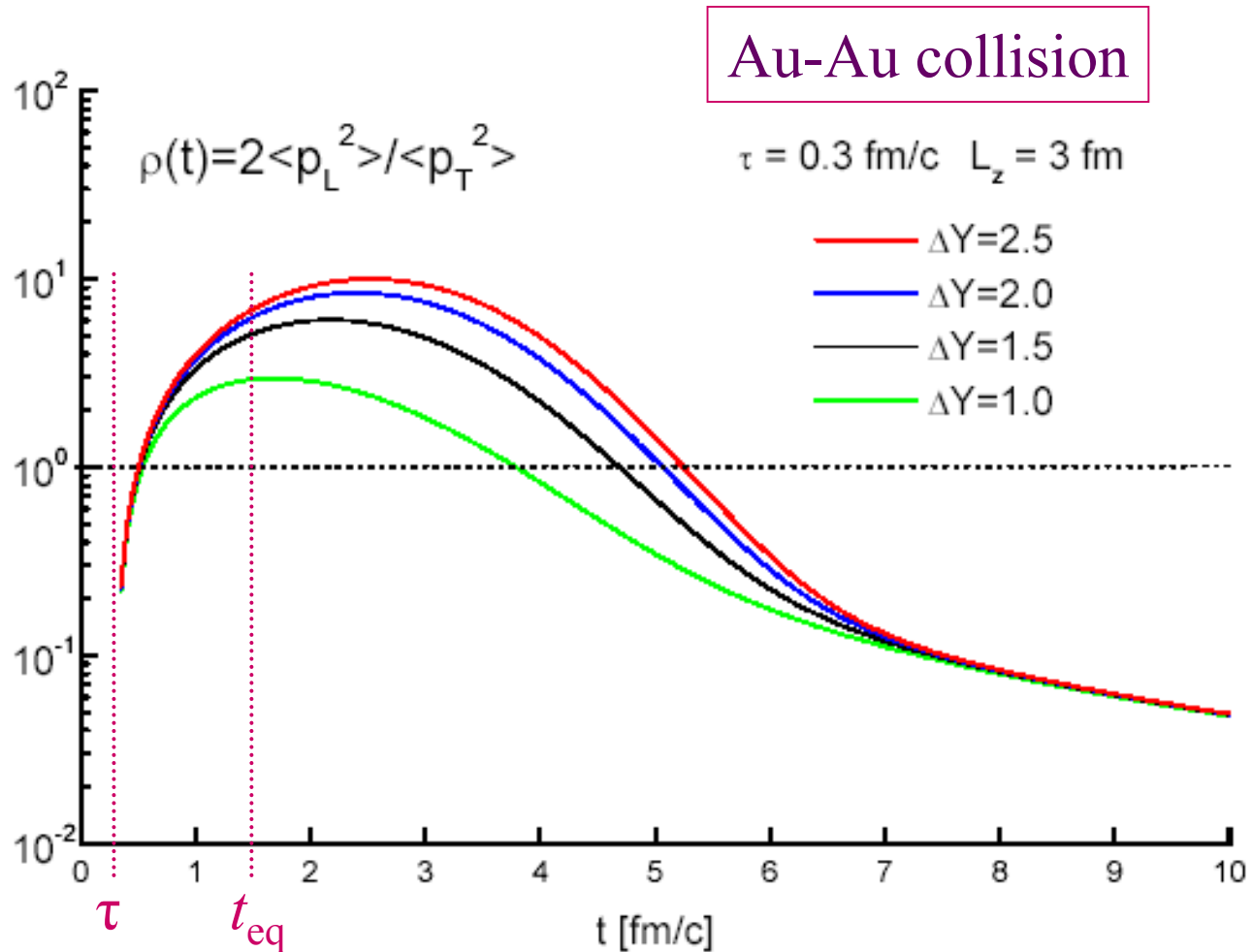


$t\sqrt{1 - v_z^2}$  - parton's proper time

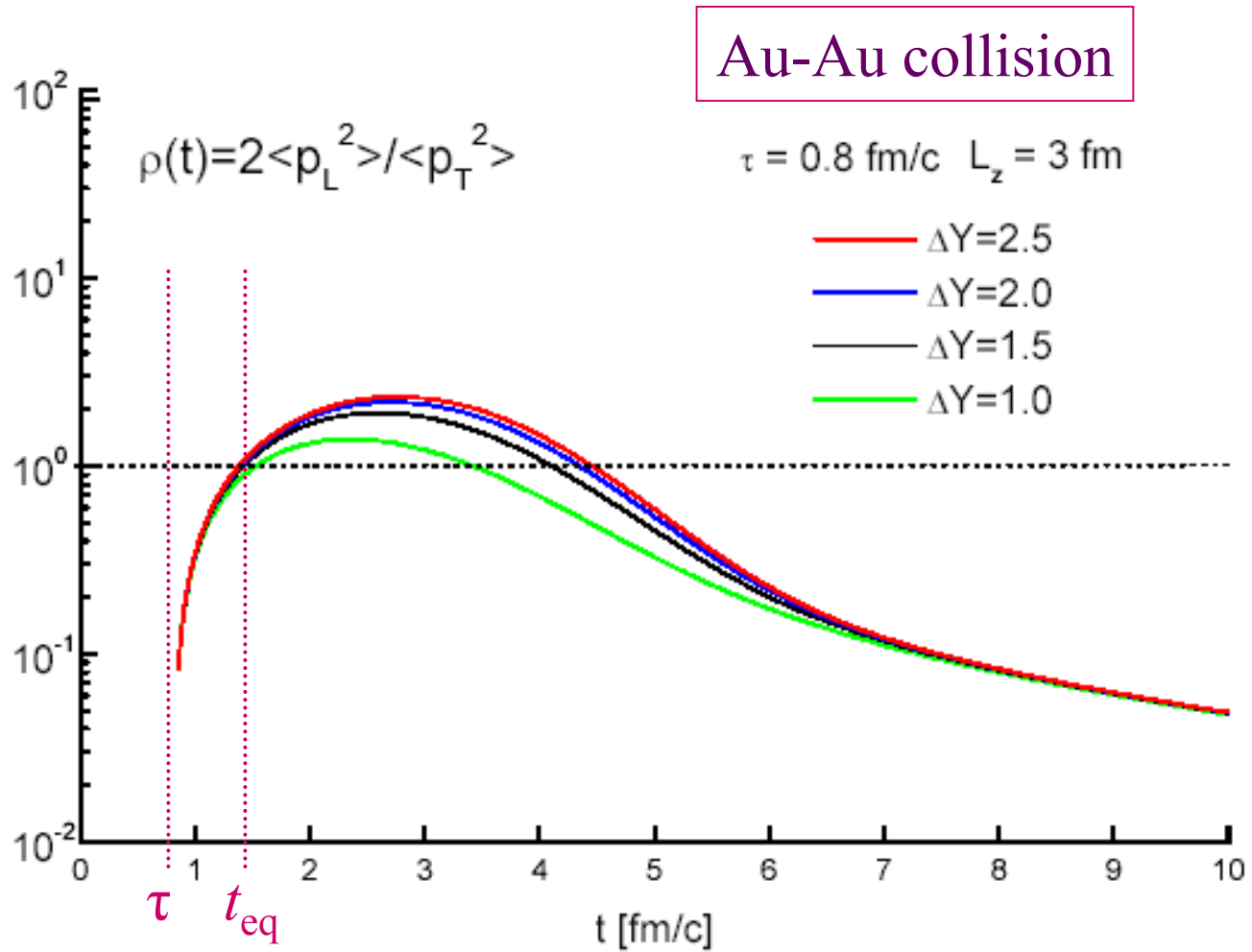
$$-Y_{\max} < Y < Y_{\max}$$

$$\text{th } Y_{\max} = \frac{1}{t} \sqrt{t^2 - \tau^2}$$

# Effect of finite formation time cont.



# Effect of finite formation time cont.



## Conclusion

**Prior to equilibrium the momentum distribution  
is prolate not oblate**

**Is the conclusion reliable?**

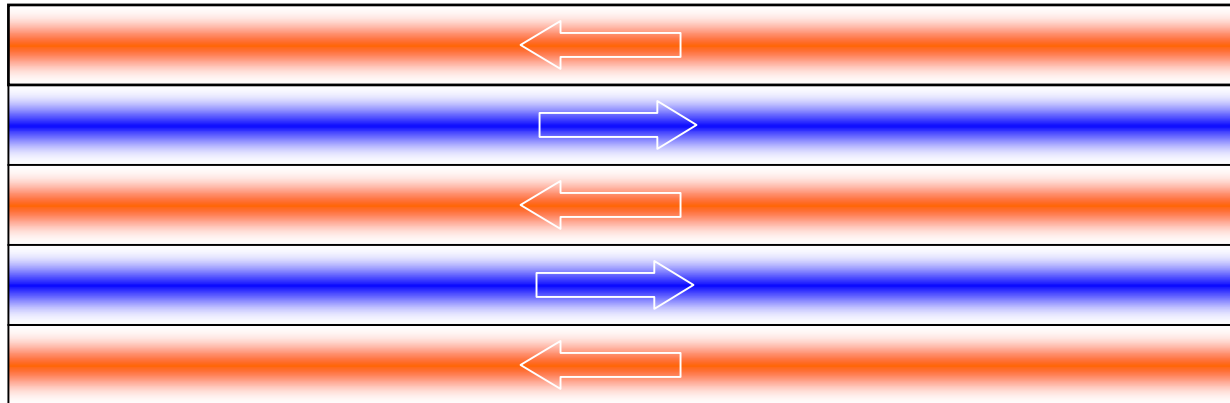
**Why the momentum configuration is important?**

# Seeds of instability

$\langle j_a^\mu(x) \rangle = 0$  but current fluctuations are finite

$$\langle j_a^\mu(x_1) j_b^\nu(x_2) \rangle = \frac{1}{8} \delta^{ab} \int \frac{d^3 p}{(2\pi)^3} \frac{p^\mu p^\nu}{E_p^2} f(\mathbf{p}) \delta^{(3)}(\mathbf{x} - \mathbf{v}t) \neq 0$$

$$x_1 = (t_1, \mathbf{x}_1), \quad x_2 = (t_2, \mathbf{x}_2), \quad x = (t_1 - t_2, \mathbf{x}_1 - \mathbf{x}_2)$$

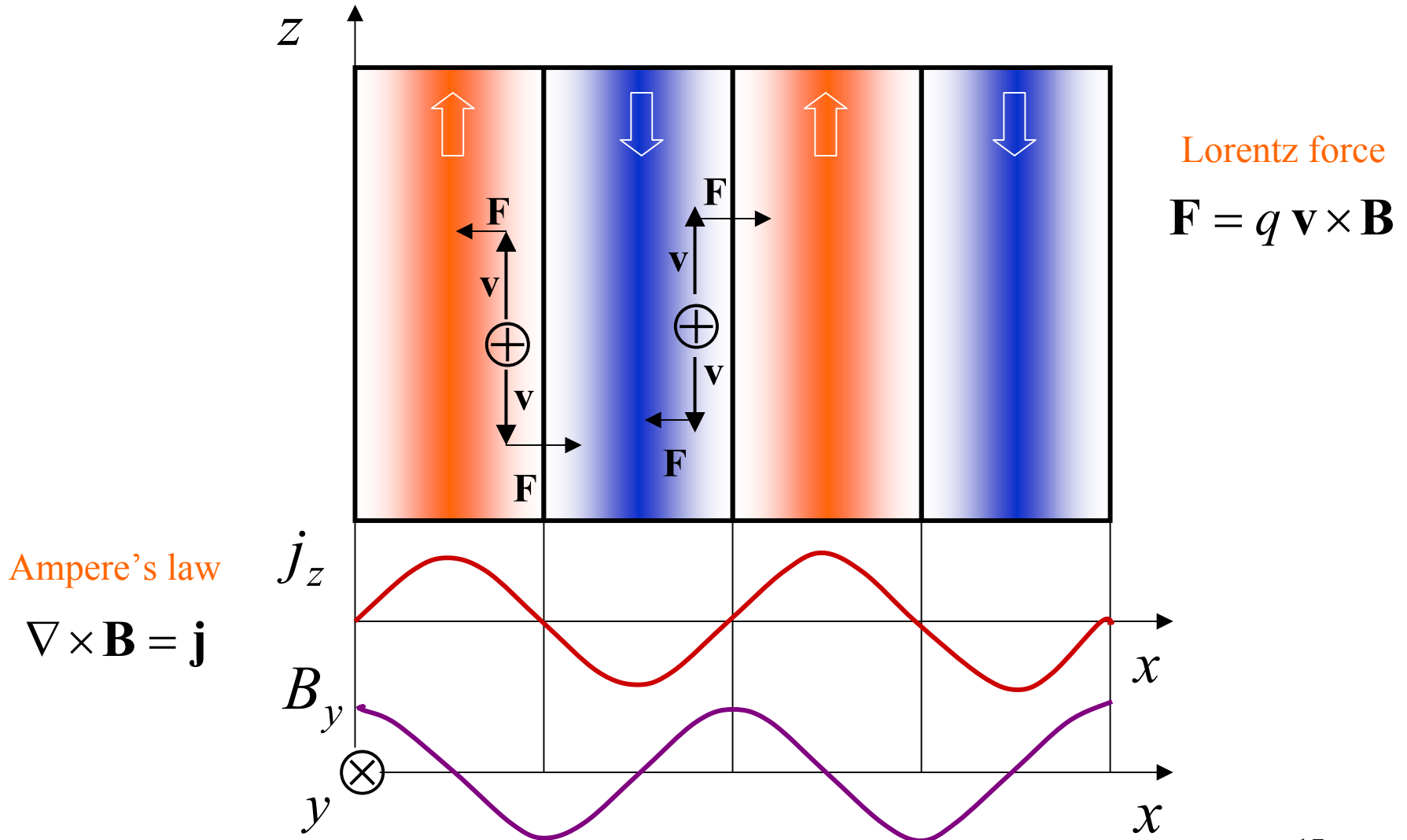


**Direction of the momentum surplus**

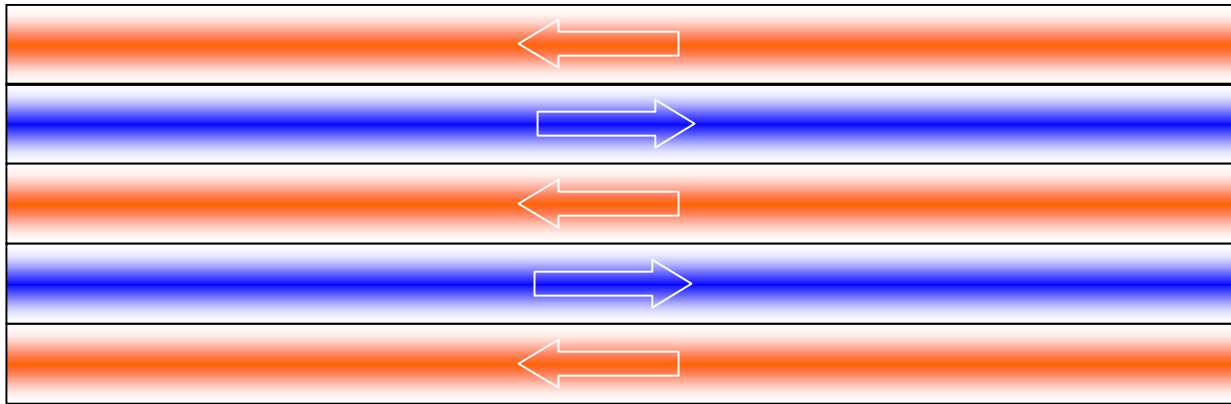




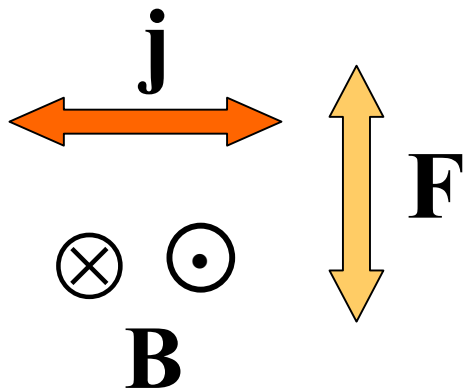
# Mechanism of filamentation



# Isotropization - particles

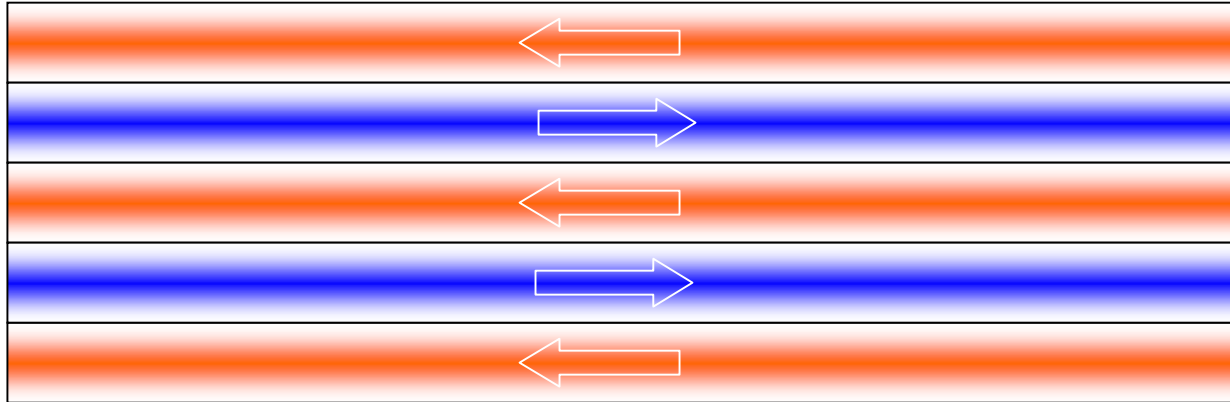


Direction of the momentum surplus

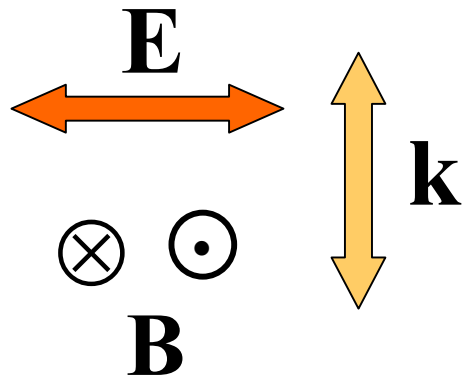


$$\Delta \mathbf{p} = \int dt \mathbf{F}$$

# Isotropization - fields



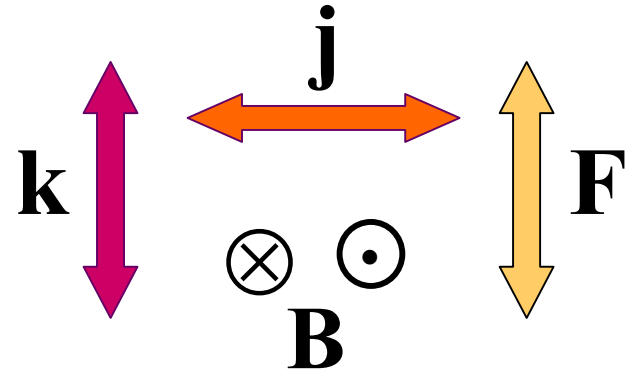
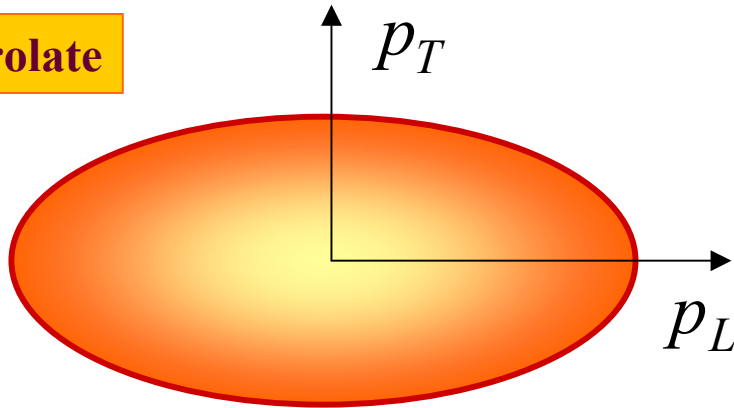
Direction of the momentum surplus



$$\mathbf{P}_{\text{fields}} \sim \mathbf{B}^a \times \mathbf{E}^a \sim \mathbf{k}$$

# Prolate vs. oblate

prolate



oblate

