

Coalescence vs. thermal model of light nuclei production in relativistic heavy-ion collisions

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with **Sylwia Bazak & Patrycja Słoń**

Background

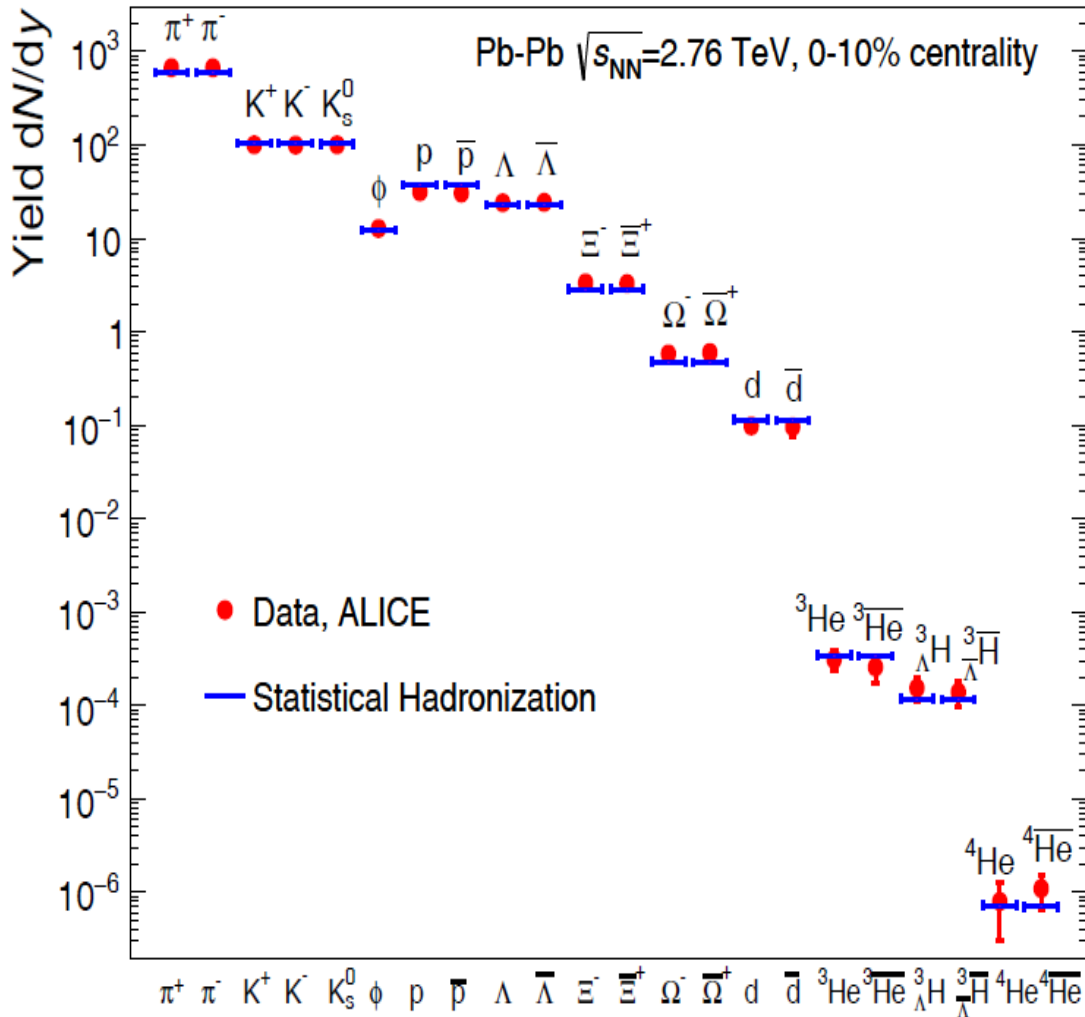
Conventional wisdom

- ▶ Production of light nuclei in relativistic heavy-ion collisions is well understood as a process of final state interactions.
- ▶ Coalescence models work well in a broad collision energy range.

News from RHIC & LHC

- ▶ Production of ${}^2\text{H}$, ${}^2\bar{\text{H}}$, ${}^3\text{H}$, ${}^3\bar{\text{H}}$, ${}^3\text{He}$, ${}^3\bar{\text{He}}$, ${}^4\text{He}$, ${}^4\bar{\text{He}}$, ${}^3_{\Lambda}\text{H}$, ${}^3_{\Lambda}\bar{\text{H}}$ is observed in midrapidity.
- ▶ Matter-antimatter symmetry is seen.
- ▶ Thermal model properly describes yields of light nuclei.

Thermal model prediction



baryonless fireball

$$\text{Yield} \sim g e^{-\frac{m}{T}}$$

$$T = 156 \text{ MeV}$$

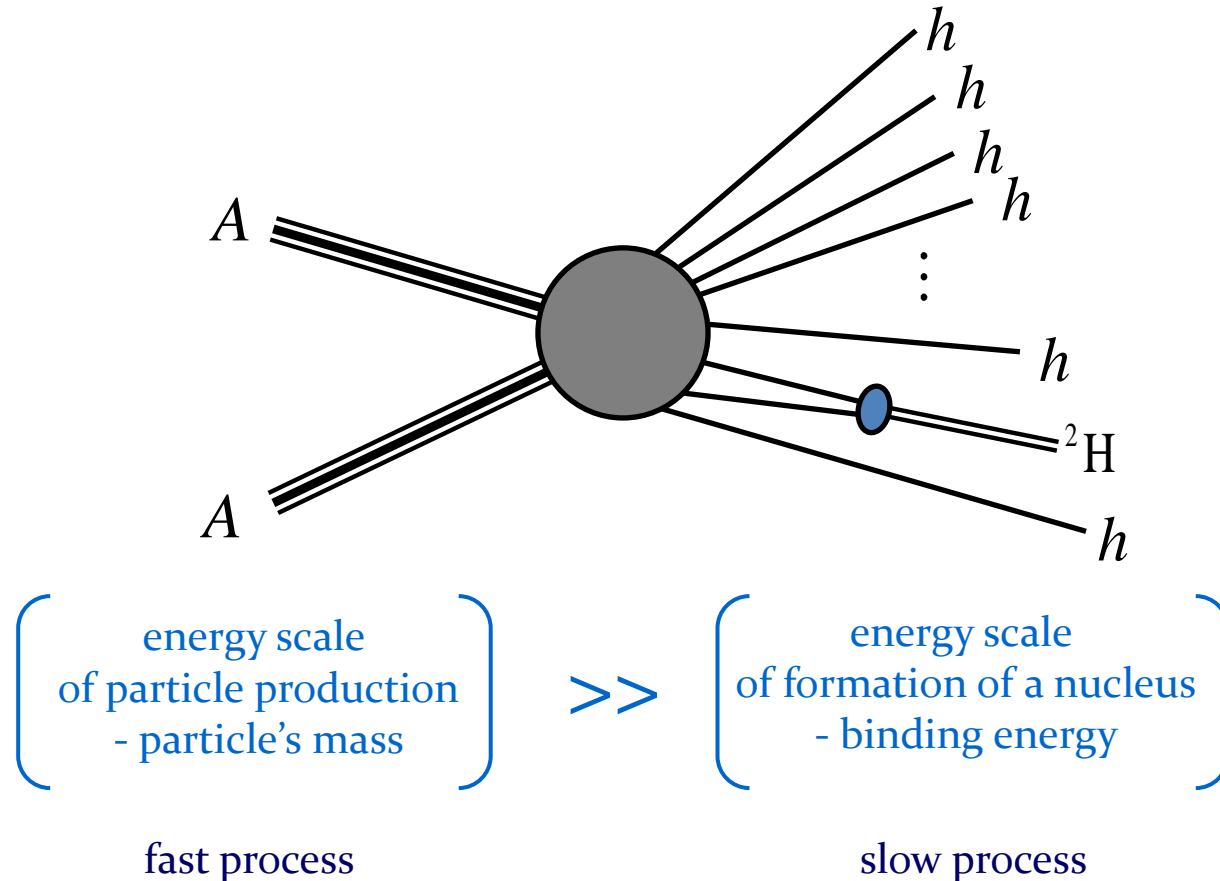
Can light nuclei exist in a fireball?

- ▶ Interparticle spacing in a hadron gas is about 1.5 fm at $T = 156$ MeV.
- ▶ Root mean square radius of a deuteron is 2.0 fm.
- ▶ Binding energy of a deuteron is $\varepsilon_B = 2.2$ MeV.
- ▶ A characteristic time of deuteron formation is $1/\varepsilon_B = 100$ fm/c.
- ▶ A hadron gas at $T = 156$ MeV is essentially a classical system.

*- Snowflakes in hell ?
- No, snowflakes from hell.*



Final state interaction – conventional approach to production of light nuclei



S.T. Butler & C.A. Pearson, Phys. Rev. **129**, 836 (1963)
A. Schwarzschild & C. Zupancic, Phys. Rev. **129**, 854 (1963)

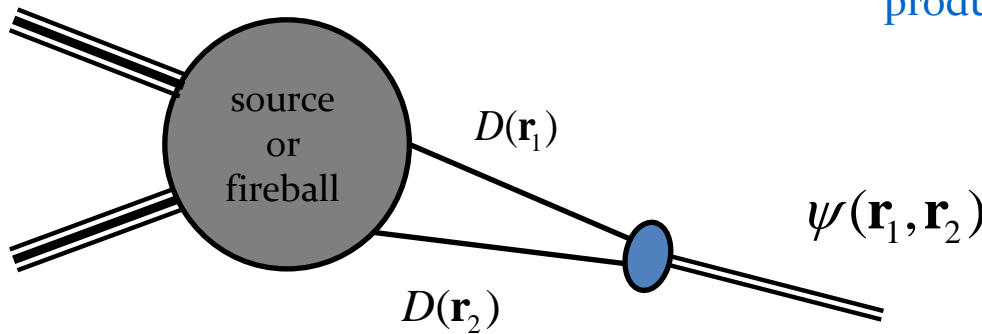
Factorization of production of nucleons and formation of a nucleus

Deuteron production cross section

$$\frac{d\sigma^D}{d^3\mathbf{P}_D} = W \frac{d\sigma^{np}}{d^3\mathbf{p}_n d^3\mathbf{p}_p} \quad \frac{1}{2}\mathbf{P}_D = \mathbf{p}_n = \mathbf{p}_p$$

deuteron formation

production of np pair



$$W = \frac{3}{4} (2\pi)^3 \int d^3\mathbf{r}_1 d^3\mathbf{r}_2 D(\mathbf{r}_1) D(\mathbf{r}_2) |\psi(\mathbf{r}_1, \mathbf{r}_2)|^2$$

H. Sato and K. Yazaki, Phys. Lett. B **98**, 153 (1981)

Deuteron formation rate vs. n-p correlation

$$\psi(\mathbf{r}_1, \mathbf{r}_2) = e^{i\mathbf{P}\cdot\mathbf{R}} \varphi(\mathbf{r}) \quad \mathbf{R} \equiv \frac{1}{2}(\mathbf{r}_1 + \mathbf{r}_2), \quad \mathbf{r} \equiv \mathbf{r}_1 - \mathbf{r}_2$$

$$W = \frac{3}{4} (2\pi)^3 \int d^3\mathbf{r} D_r(\mathbf{r}) |\varphi(\mathbf{r})|^2$$

$$D_r(\mathbf{r}) \equiv \int d^3\mathbf{R} D\left(\mathbf{R} - \frac{1}{2}\mathbf{r}\right) D\left(\mathbf{R} + \frac{1}{2}\mathbf{r}\right)$$

n-p – correlation function

$$C(\mathbf{q}) = \int d^3\mathbf{r} D_r(\mathbf{r}) |\varphi_{\mathbf{q}}(\mathbf{r})|^2$$

$\varphi(\mathbf{r})$ – wave function of a bound state

$\varphi_{\mathbf{q}}(\mathbf{r})$ – wave function of a scattering state

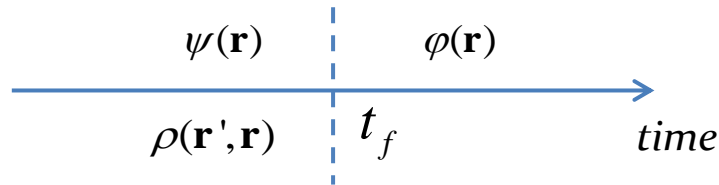
If emission time included

$$R_s \rightarrow \sqrt{R_s^2 + v^2 \tau^2}$$

St. Mrówczyński, Phys. Lett. B 277, 43 (1992)

Quantum-mechanical meaning of the formation rate formula

Sudden approximation



Transition matrix element

$$W = \left| \int d^3\mathbf{r} \psi^*(\mathbf{r}) \phi(\mathbf{r}) \right|^2 = \int d^3\mathbf{r} d^3\mathbf{r}' \phi^*(\mathbf{r}') \underbrace{\psi(\mathbf{r}') \psi^*(\mathbf{r})}_{\rho(\mathbf{r}', \mathbf{r})} \phi(\mathbf{r})$$

density matrix

$$W = \int d^3\mathbf{r} d^3\mathbf{r}' \phi^*(\mathbf{r}') \rho(\mathbf{r}', \mathbf{r}) \phi(\mathbf{r})$$

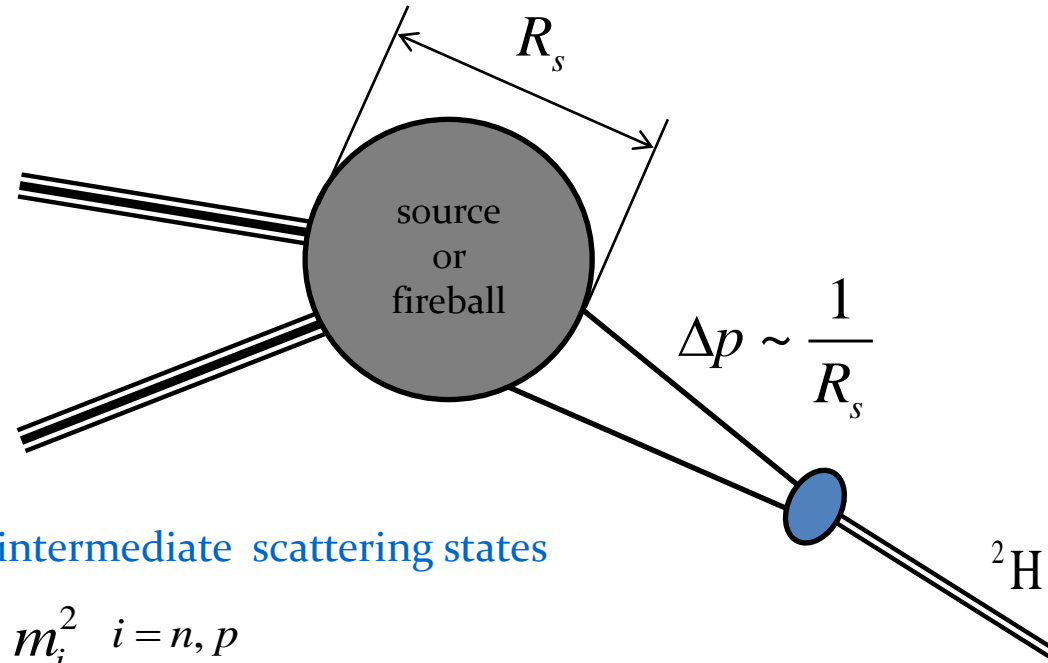
If density matrix is diagonal (random phase approximation)

$$\rho(\mathbf{r}', \mathbf{r}) = D(\mathbf{r}) \delta^{(3)}(\mathbf{r}' - \mathbf{r})$$

\Rightarrow

$$W = \int d^3\mathbf{r} D(\mathbf{r}) |\phi(\mathbf{r})|^2$$

Energy-momentum conservation



Nucleons are intermediate scattering states

$$E_i^2 - \mathbf{p}_i^2 \neq m_i^2 \quad i = n, p$$

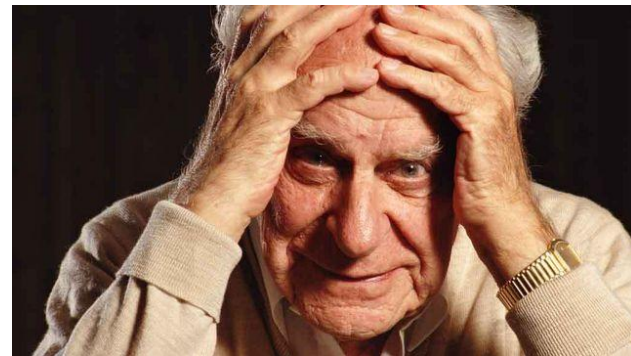
Energy-momentum conservation

$$\begin{cases} \mathbf{p}_p + \mathbf{p}_n = \mathbf{p}_D \\ E_p + E_n = E_D \end{cases}$$

St. Mrówczyński, J. Phys. G **11**, 1087 (1987)

Thermal vs. coalescence model


- ▶ The two models usually give quantitatively similar predictions.
- ▶ How to falsify one of the models experimentally?



Karl Popper 1902-1994

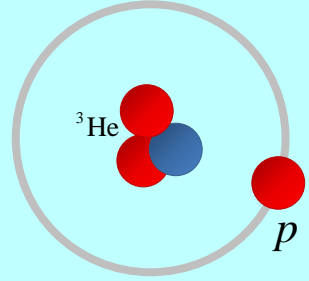
The first idea: ${}^4\text{He}$ vs. ${}^4\text{Li}$

${}^4\text{He}$



$r_{\text{RMS}} = 1.68 \text{ fm}$
 $\varepsilon_B = 28.3 \text{ MeV}$
 $m = 3727.4 \text{ MeV}$
 $s = 0$

${}^4\text{Li}$



${}^4\text{Li} \rightarrow {}^3\text{He} + p$
 $\Gamma = 6 \text{ MeV}$
 $m = m_{{}^3\text{He}} + m_p + 4.1 \text{ MeV}$
 $m = 3749.7 \text{ MeV}$
 $s = 2$

▶ Thermal model $\frac{\text{Yield}({}^4\text{Li})}{\text{Yield}({}^4\text{He})} = \frac{2S_{\text{Li}} + 1}{2S_{\text{He}} + 1} = 5$

▶ Coalescence model $\frac{\text{Yield}({}^4\text{Li})}{\text{Yield}({}^4\text{He})} = \frac{W_{\text{Li}}}{W_{\text{He}}}$

Formation rates of ${}^4\text{He}$ & ${}^4\text{Li}$

$$W = g_s g_I (2\pi)^9 \int d^3\mathbf{r}_1 d^3\mathbf{r}_2 \int d^3\mathbf{r}_3 d^3\mathbf{r}_4 D(\mathbf{r}_1) D(\mathbf{r}_2) D(\mathbf{r}_3) D(\mathbf{r}_4) |\psi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \mathbf{r}_4)|^2$$

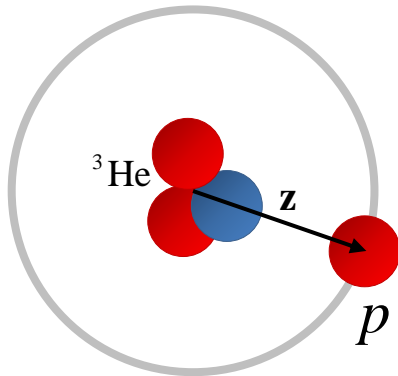
▶ ${}^4\text{He}$



$$\mathbf{r}_{ij} \equiv \mathbf{r}_i - \mathbf{r}_j$$

$$|\psi_{\text{He}}(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \mathbf{r}_4)|^2 \sim \exp\left[-\alpha(\mathbf{r}_{12}^2 + \mathbf{r}_{13}^2 + \mathbf{r}_{14}^2 + \mathbf{r}_{23}^2 + \mathbf{r}_{24}^2 + \mathbf{r}_{34}^2)\right]$$

▶ ${}^4\text{Li}$



J. C. Bergstrom, Nucl. Phys. A **327**, 458 (1979)

$$\mathbf{z} \equiv \mathbf{r}_4 - \frac{1}{3}(\mathbf{r}_1 + \mathbf{r}_2 + \mathbf{r}_3)$$

$$|\psi_{\text{Li}}(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \mathbf{r}_4)|^2 \sim \exp\left[-\beta(\mathbf{r}_{12}^2 + \mathbf{r}_{13}^2 + \mathbf{r}_{23}^2)\right] \mathbf{z}^4 \exp(-\gamma \mathbf{z}^2) |Y_{lm}(\Omega_{\mathbf{z}})|^2$$

Formation rates of ${}^4\text{He}$ & ${}^4\text{Li}$

$$W = g_S g_I (2\pi)^9 \int d^3\mathbf{r}_1 d^3\mathbf{r}_2 \int d^3\mathbf{r}_3 d^3\mathbf{r}_4 D(\mathbf{r}_1) D(\mathbf{r}_2) D(\mathbf{r}_3) D(\mathbf{r}_4) |\psi(\mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \mathbf{r}_4)|^2$$

Source function

$$D(\mathbf{r}_i) = \frac{1}{(2\pi R_s^2)^{3/2}} \exp\left(-\frac{\mathbf{r}_i^2}{2R_s^2}\right) \quad i = 1, 2, 3, 4$$

If emission time included

$$R_s \rightarrow \sqrt{R_s^2 + v^2 \tau^2}$$

Jacobi variables

$$\left\{ \begin{array}{l} \mathbf{R} \equiv \frac{1}{4}(\mathbf{r}_1 + \mathbf{r}_2 + \mathbf{r}_3 + \mathbf{r}_4) \\ \mathbf{x} \equiv \mathbf{r}_2 - \mathbf{r}_1 \\ \mathbf{y} \equiv \mathbf{r}_3 - \frac{1}{2}(\mathbf{r}_1 + \mathbf{r}_2) \\ \mathbf{z} \equiv \mathbf{r}_4 - \frac{1}{3}(\mathbf{r}_1 + \mathbf{r}_2 + \mathbf{r}_3) \end{array} \right.$$

$$\blacktriangleright \quad \mathbf{r}_1^2 + \mathbf{r}_2^2 + \mathbf{r}_3^2 + \mathbf{r}_4^2 = 4\mathbf{R}^2 + \frac{1}{2}\mathbf{x}^2 + \frac{2}{3}\mathbf{y}^2 + \frac{3}{4}\mathbf{z}^2$$

$$\blacktriangleright \quad \mathbf{r}_{12}^2 + \mathbf{r}_{13}^2 + \mathbf{r}_{14}^2 + \mathbf{r}_{23}^2 + \mathbf{r}_{24}^2 + \mathbf{r}_{34}^2 = 2\mathbf{x}^2 + \frac{8}{3}\mathbf{y}^2 + 3\mathbf{z}^2$$

$$\mathbf{r}_{ij} \equiv \mathbf{r}_i - \mathbf{r}_j$$

Fully analytic calculations
are possible!

Formation rates of ${}^4\text{He}$ & ${}^4\text{Li}$

$$\triangleright W_{\text{He}} = \frac{\pi^{9/2}}{2^{9/2}} \frac{1}{\left(R_s^2 + R_\alpha^2\right)^{9/2}}$$

$$\triangleright W_{\text{Li}} = \frac{3\pi^{9/2}}{2^{11/2}} \begin{pmatrix} 5 \\ 2 \\ 1 \end{pmatrix} \frac{R_s^4}{\left(R_s^2 + \frac{1}{2}R_c^2\right)^3 \left(R_s^2 + \frac{4}{7}R_{\text{Li}}^2 - \frac{3}{7}R_c^2\right)^{7/2}} \begin{pmatrix} l=1 \\ l=2 \end{pmatrix}$$

Since ${}^4\text{Li}$ is $J^P = 2^-$ then $l=1$.

R_s – root mean square radius of the source

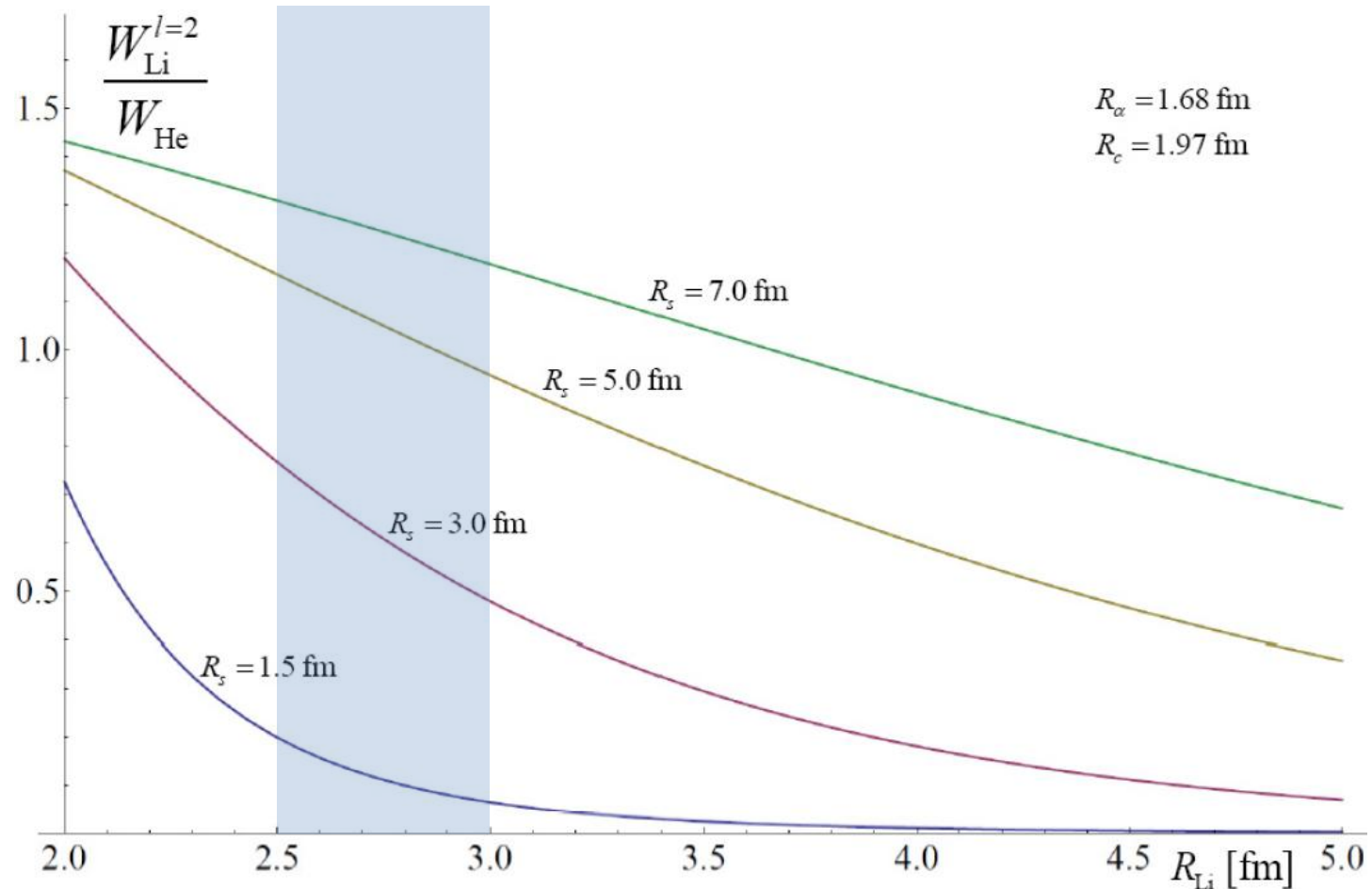
R_α – root mean square radius of ${}^4\text{He}$

R_{Li} – root mean square radius of ${}^4\text{Li}$

R_c – root mean square radius of ${}^3\text{He}$ cluster in ${}^4\text{Li}$

Ratio of yields of ${}^4\text{Li}$ to ${}^4\text{He}$

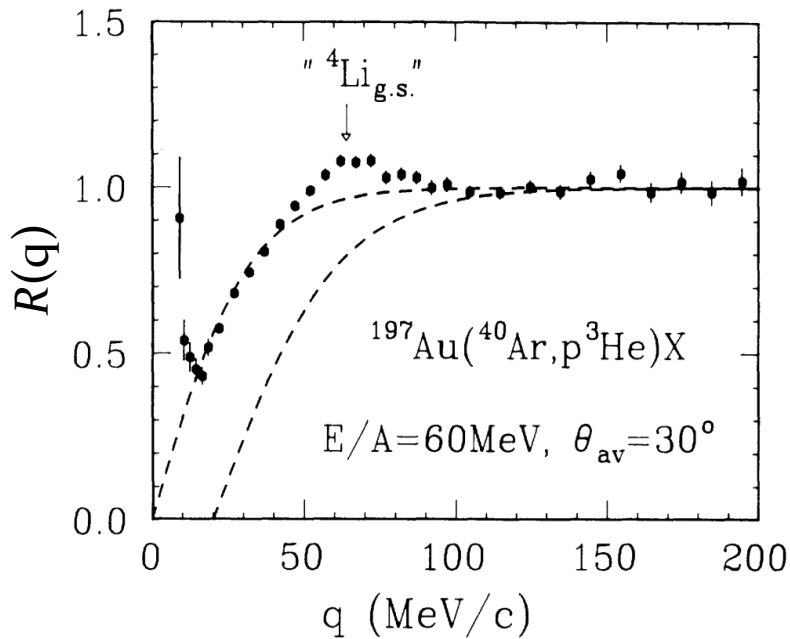
In the thermal model the ratio equals 5.



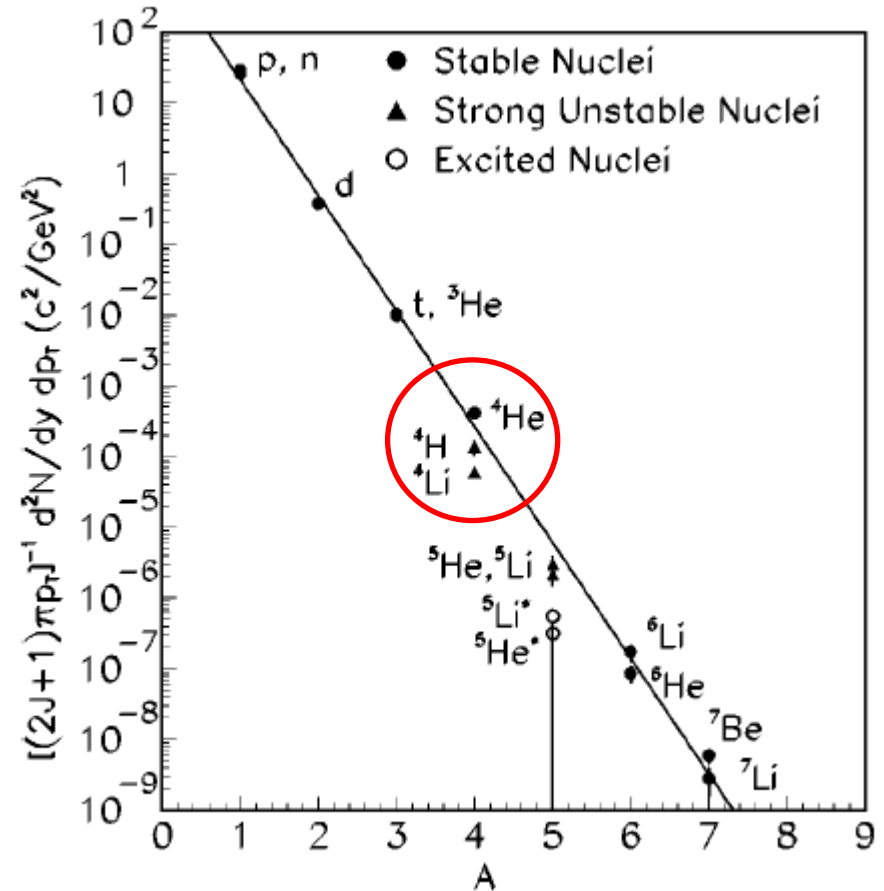
How to observe ${}^4\text{Li}$?

Measurement of the correlation function of ${}^3\text{He}$ - p is needed

${}^{197}\text{Au}+{}^{197}\text{Pt}$ @ AGS

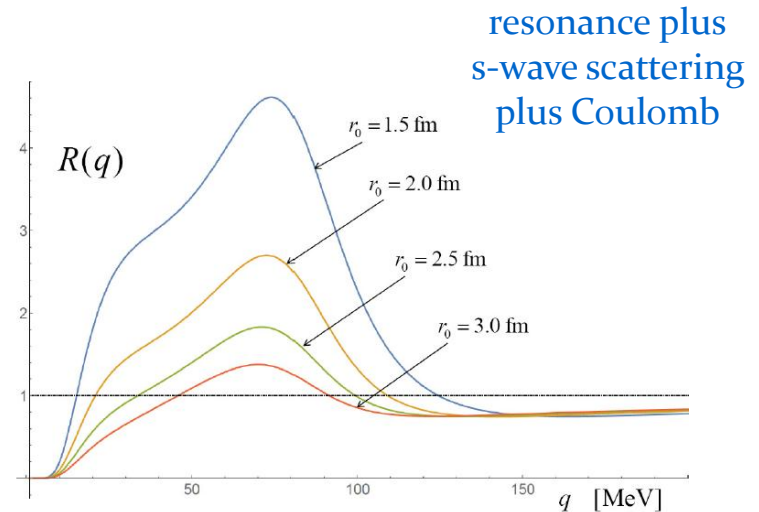
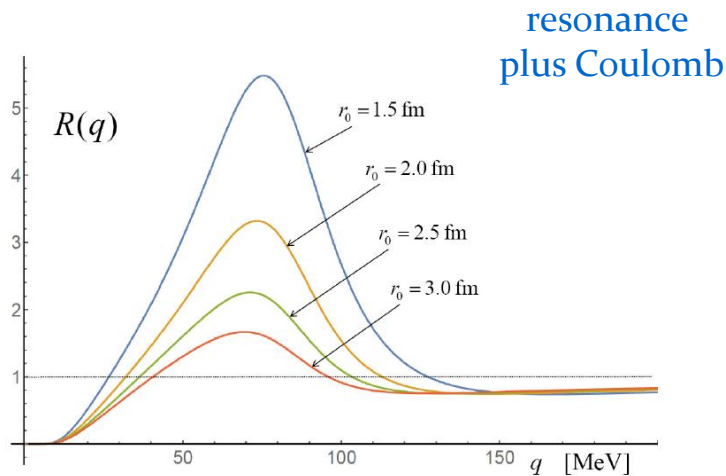
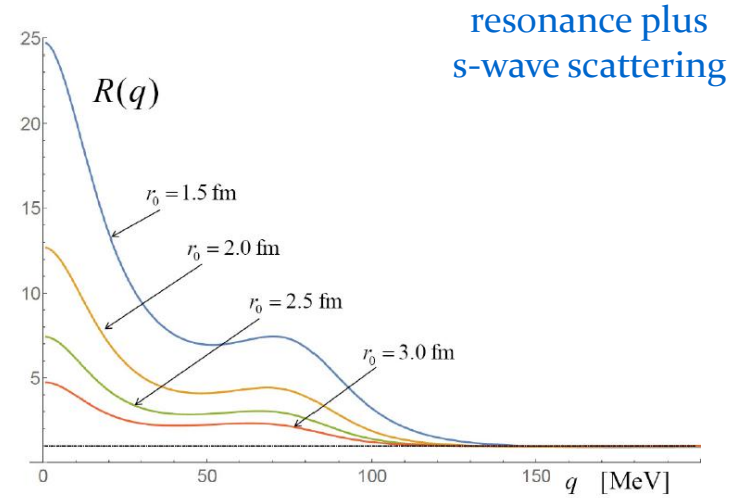
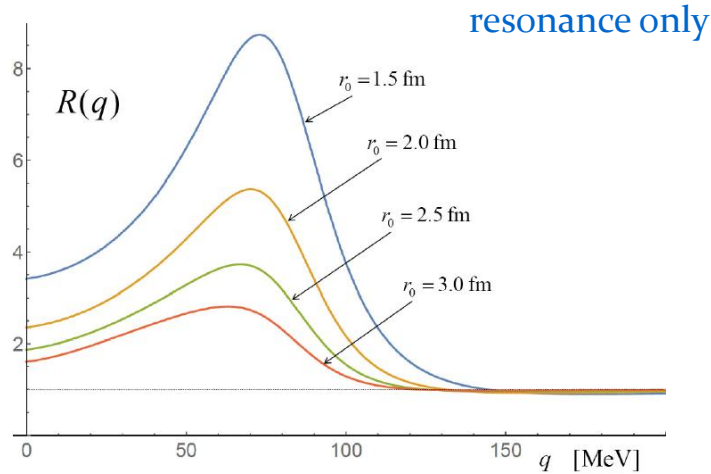


J. Pochodzala et al. Phys. Rev. C **35**, 1695 (1987)



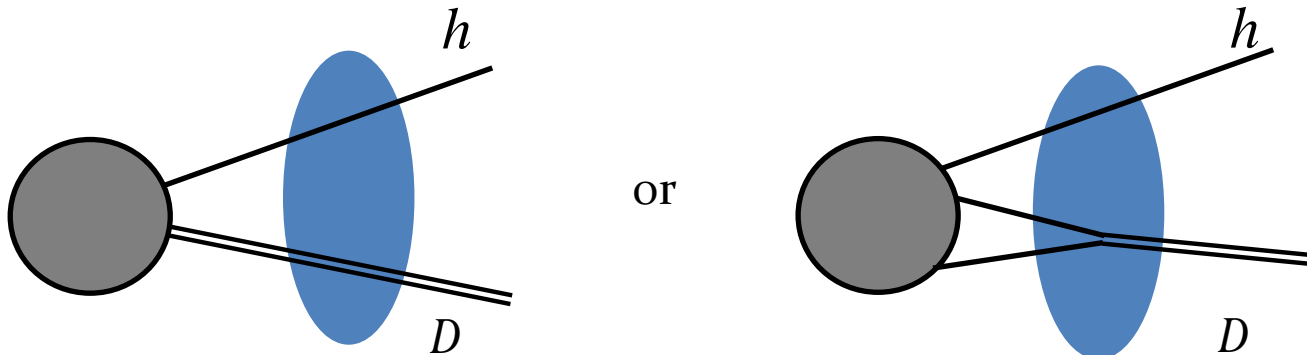
T. A. Armstrong et al. Phys. Rev. C **65**, 014906 (2001)

Correlation function p - ^3He



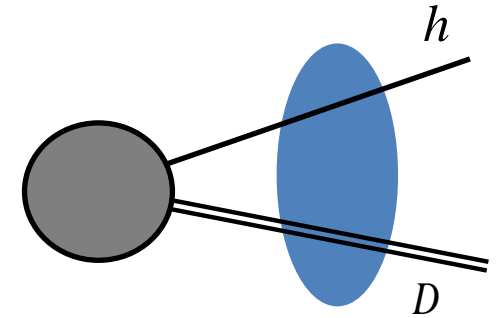
The second idea: h - D correlations

- ▶ Hadron-deuteron correlations carry information about a source of deuterons.
- ▶ A measurement of p - D & p - p correlation functions is suggested to falsify the thermal or coalescence model.



Hadron-deuteron correlation function

1) Deuteron is treated as an elementary particle



Experimental definition

$$\frac{dN_{hD}}{d\mathbf{p}_h d\mathbf{p}_D} = R(\mathbf{p}_h, \mathbf{p}_D) \frac{dN_h}{d\mathbf{p}_h} \frac{dN_D}{d\mathbf{p}_D}$$

Theoretical formula

$$R(\mathbf{p}_h, \mathbf{p}_D) = \int d^3r_h d^3r_D D(\mathbf{r}_h) D(\mathbf{r}_D) |\psi(\mathbf{r}_h, \mathbf{r}_D)|^2$$

distribution
of emission points

h - D wave function

S.E. Koonin, Phys. Lett. B **70**, 43 (1977)

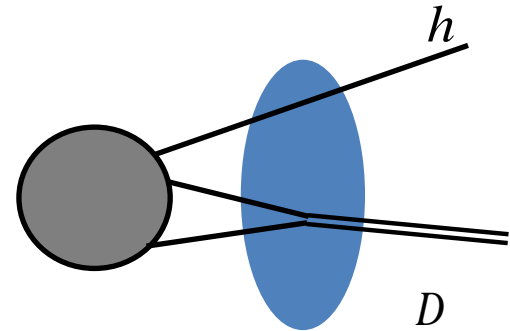
R. Lednicky and V.L. Lyuboshitz, Yad. Fiz. **35**, 1316 (1982)

Hadron-deuteron correlation function

2) Deuteron is treated as a bound state of neutron and proton

Experimental definition

$$\frac{dN_{hD}}{d\mathbf{p}_h d\mathbf{p}_D} = R(\mathbf{p}_h, \mathbf{p}_D) W \frac{dN_h}{d\mathbf{p}_h} \frac{dN_n}{d\mathbf{p}_n} \frac{dN_p}{d\mathbf{p}_p}$$



Theoretical formula

$$R(\mathbf{p}_h, \mathbf{p}_D) = \frac{1}{W} \int d^3 r_h d^3 r_n d^3 r_p D(\mathbf{r}_h) D(\mathbf{r}_n) D(\mathbf{r}_p) |\psi_{hD}(\mathbf{r}_h, \mathbf{r}_n, \mathbf{r}_p)|^2$$

Deuteron formation rate

$$\frac{dN_D}{d\mathbf{p}_D} = W \frac{dN_n}{d\mathbf{p}_n} \frac{dN_p}{d\mathbf{p}_p} \quad \frac{1}{2} \mathbf{p}_D = \mathbf{p}_n = \mathbf{p}_p$$

$$W = \frac{3}{4} (2\pi)^3 \int d^3 \mathbf{r}_n d^3 \mathbf{r}_p D(\mathbf{r}_n) D(\mathbf{r}_p) |\psi_D(\mathbf{r}_n, \mathbf{r}_p)|^2 = \frac{3}{4} (2\pi)^3 \int d^3 r_{np} D_r(\mathbf{r}_{np}) |\phi_D(\mathbf{r}_{np})|^2$$

spin factor

$$\psi_D(\mathbf{r}_n, \mathbf{r}_p) = e^{i\mathbf{P}\mathbf{R}} \phi_D(\mathbf{r}_{np})$$

Thermal vs. coalescence model

Thermal model

$$R(\mathbf{q}) = \int d^3 r_{hD} D_r(\mathbf{r}_{hD}) |\phi_{\mathbf{q}}(\mathbf{r}_{hD})|^2$$

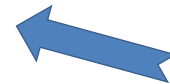
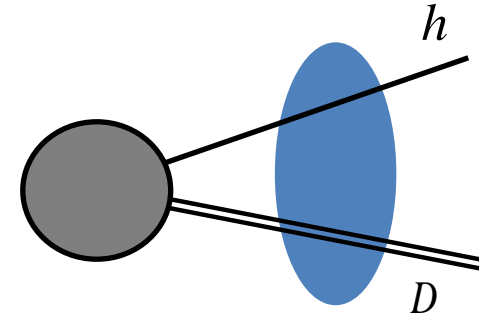


$$D_r(\mathbf{r}) = \left(\frac{1}{4\pi R^2} \right)^{3/2} \exp\left(-\frac{\mathbf{r}^2}{4R^2} \right)$$

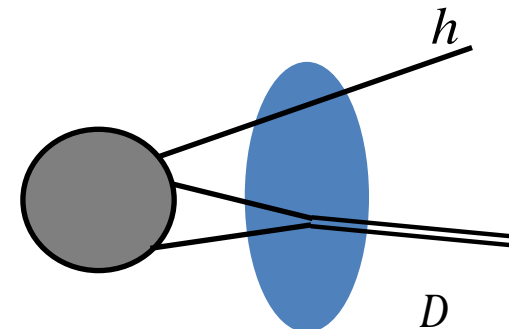
$$D_{3r}(\mathbf{r}) = \left(\frac{1}{3\pi R^2} \right)^{3/2} \exp\left(-\frac{\mathbf{r}^2}{3R^2} \right)$$

Coalescence model

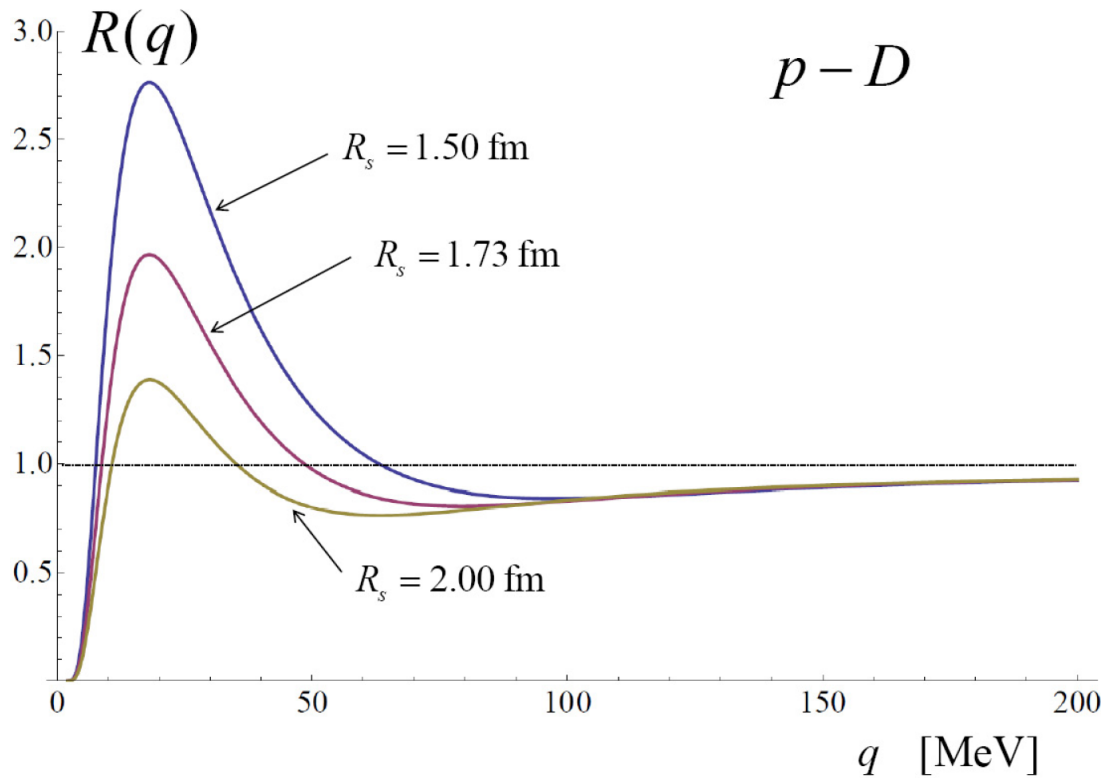
$$R(\mathbf{q}) = \int d^3 r_{hD} D_{3r}(\mathbf{r}_{hD}) |\phi_{\mathbf{q}}(\mathbf{r}_{hD})|^2$$



$$D(\mathbf{r}) = \left(\frac{1}{2\pi R^2} \right)^{3/2} \exp\left(-\frac{\mathbf{r}^2}{2R^2} \right)$$



p-D correlation functions



$$R(q) = \frac{1}{3} R_{1/2}(q) + \frac{2}{3} R_{3/2}(q)$$

$$a_{1/2} = 4.0 \text{ fm}$$

$$a_{3/2} = 11.0 \text{ fm}$$

$$2.00 = \sqrt{\frac{4}{3}} 1.73 = \frac{4}{3} 1.50$$

R_s from *p-D* correlation function vs. *R_s* from *p-p* correlation function

Conclusions

${}^4\text{He}$ vs. ${}^4\text{Li}$



The thermal and coalescence models give different predictions on the ratio of yields of ${}^4\text{Li}$ to ${}^4\text{He}$.



In the thermal model the ratio of yields is independent of collision centrality.



In the coalescence model the ratio is maximal for central collisions and rapidly decreases when one goes to peripheral collisions.



Since ${}^4\text{Li}$ can be observed through the correlation function of ${}^3\text{He}$ - p , the correlation needs to be measured.

h - D correlations



Hadron-deuteron correlations carry information about source of deuterons.



Measurement of h - D & h - p correlation function can tell us whether deuterons are directly emitted from a fireball like protons or deuterons are formed due to final state interactions.



p - D correlation functions show a sufficient sensitivity to a size of particle source to falsify the thermal or coalescence model.