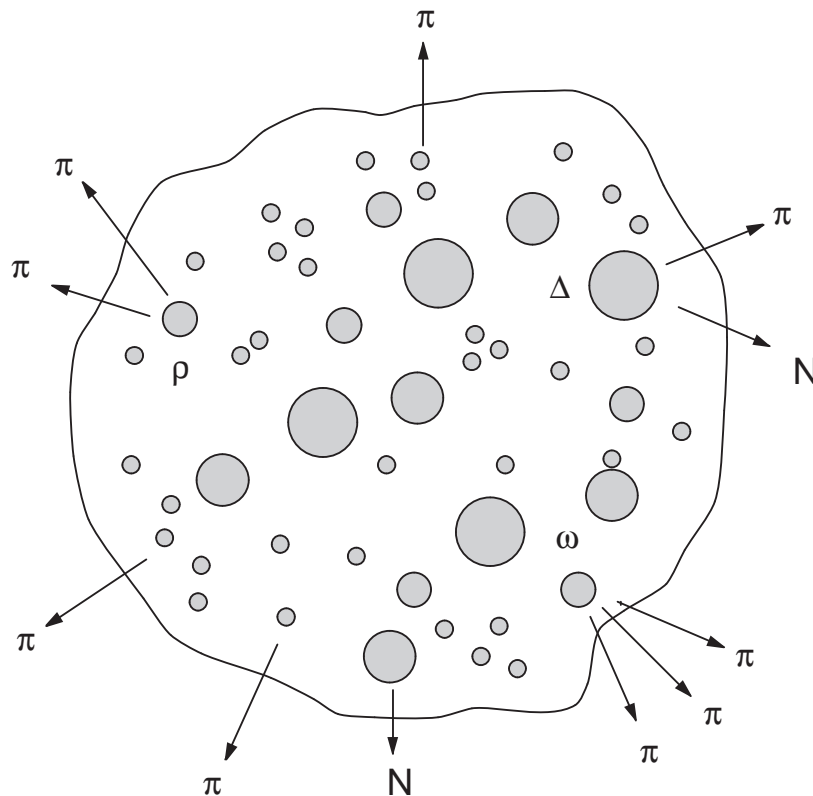


Termalny opis danych z RHIC'a

WB + W. Florkowski + Anna Baran



100 lat, Panie Profesorze!

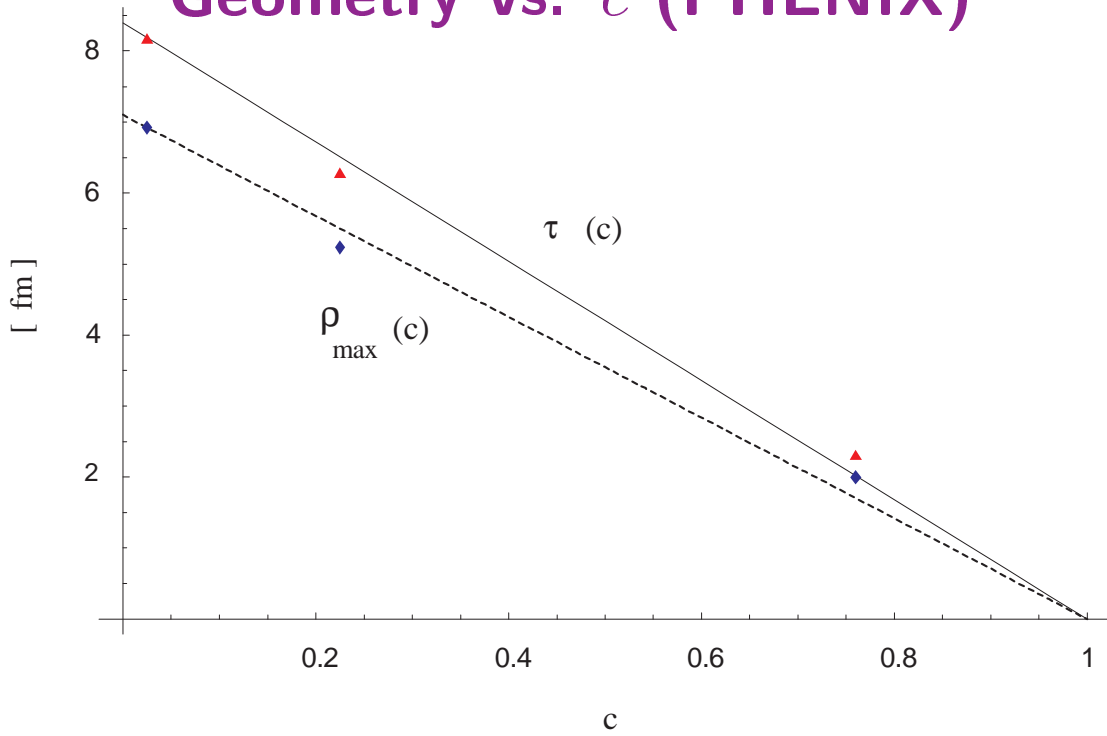
Multiplicity vs. energy

[from A. Budzanowski, *By small steps towards "The Beginning". What have we learned from first results of the PHOBOS detector at RHIC?*, Acta. Phys. Pol. **B 33** (2002) 33]

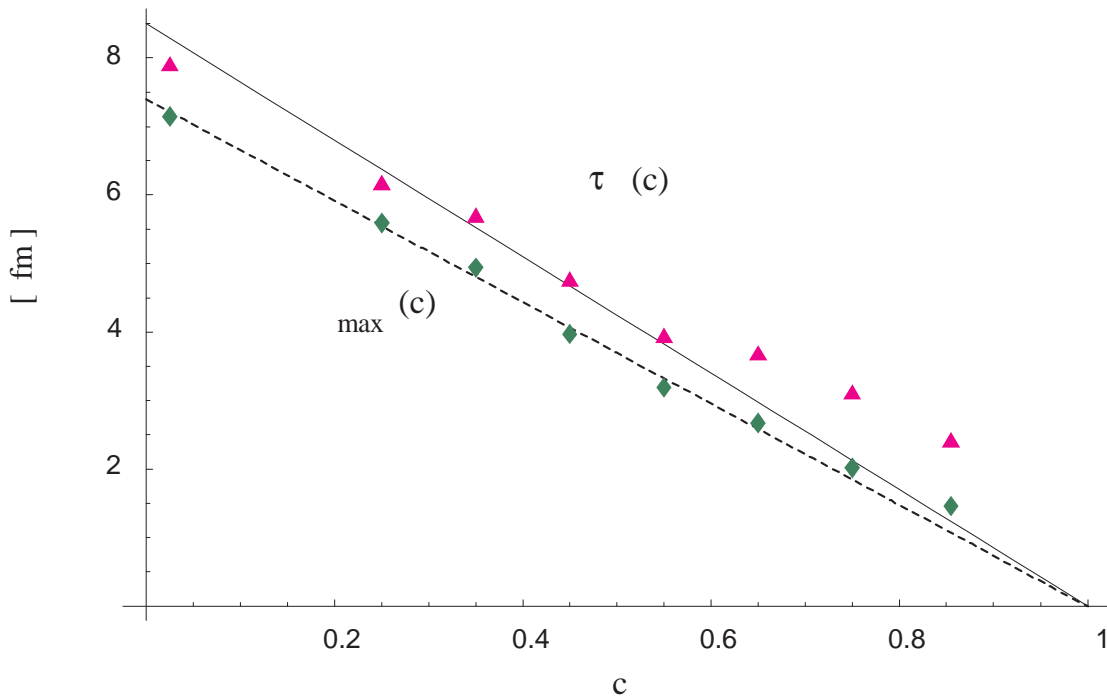
Ratios @ 200 GeV

| | Model | Experiment |
|---|-------------|---|
| Fitted thermal parameters | | |
| T [MeV] | 160 ± 5 | |
| μ_B [MeV] | 26 ± 4 | |
| μ_S [MeV] | 5 | |
| μ_I [MeV] | -1 | |
| χ^2/n | 1.5 | |
| Ratios used in the thermal analysis | | |
| π^-/π^+ | 1.015 | $1.025 \pm 0.006 \pm 0.018$ (0-12%) $1.02 \pm 0.02 \pm 0.10$ (0-5%) |
| K^-/K^+ | 0.95 | $0.95 \pm 0.03 \pm 0.03$ (0-12%) $0.92 \pm 0.03 \pm 0.10$ (0-5%) |
| \bar{p}/p | 0.74 | $0.73 \pm 0.02 \pm 0.03$ (0-12%) $0.70 \pm 0.04 \pm 0.10$ (0-5%) 0.78 ± 0.05 (0-5%) |
| \bar{p}/π^- | 0.089 | 0.083 ± 0.015 (0-5%) |
| K^-/π^- | 0.174 | 0.156 ± 0.020 (0-5%) |
| $\Omega/h^- \times 10^3$ | 0.841 | $0.887 \pm 0.111 \pm 0.133$ (0-10%) |
| $\bar{\Omega}/h^- \times 10^3$ | 0.740 | $0.935 \pm 0.105 \pm 0.140$ (0-10%) |
| $K^*(892)/\pi^-$ | 0.055 | $0.030 \pm 0.004 \pm 0.007$ (0-20%) |
| $K^*(892)/K^-$ | 0.32 | $0.19 \pm 0.02 \pm 0.05$ (0-20%) |
| $\phi/K^*(892)$ | 0.44 | $0.63 \pm 0.10 \pm 0.16$ (0-20%) |
| Preliminary STAR data on ρ and f_0 | | |
| ρ^0/π^- | 0.12 | $0.20 \pm 0.03 \pm 0.06$ (40-80%) |
| $f_0(980)/\pi^-$ | 0.01 | $0.05 \pm 0.03 \pm 0.03$ (40-80%) |

Geometry vs. c (PHENIX)



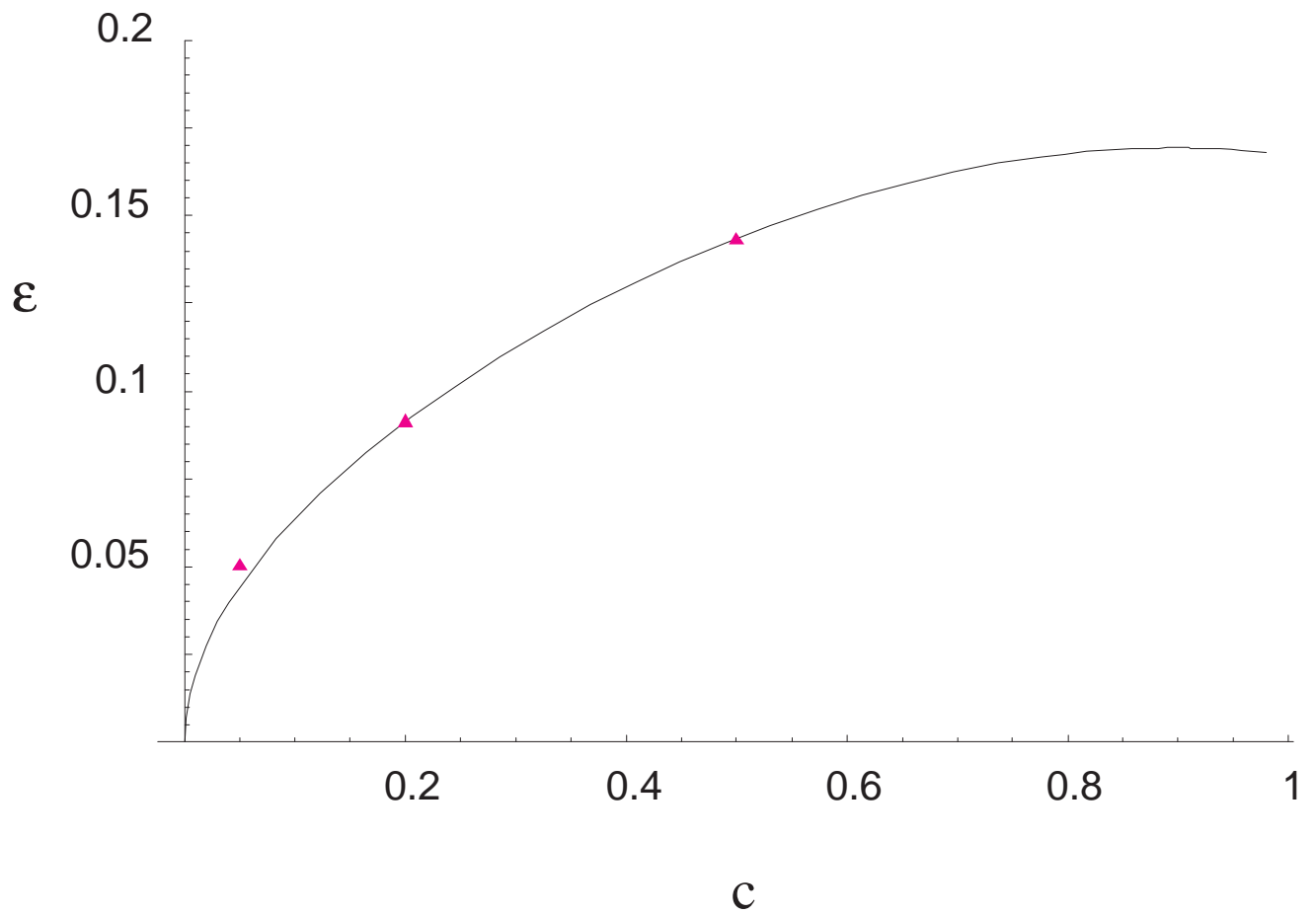
@ 130 GeV, $T = 165$ MeV and $\mu_B = 41$ MeV



@ 200 GeV, $T = 165$ MeV and $\mu_B = 26$ MeV

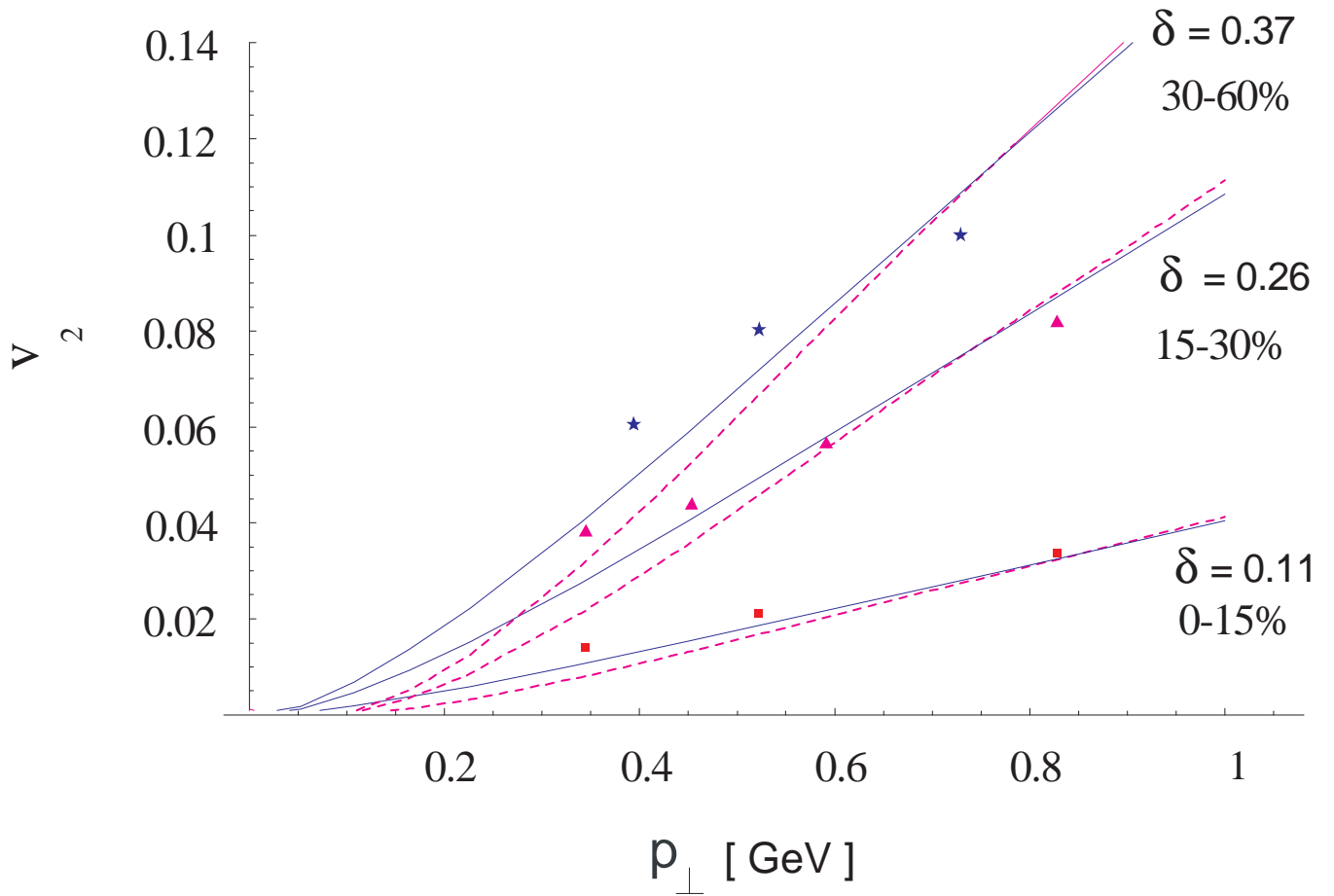
Deformation vs. c

$$\varepsilon = \frac{R_y^2 - R_x^2}{R_y^2 + R_x^2}$$

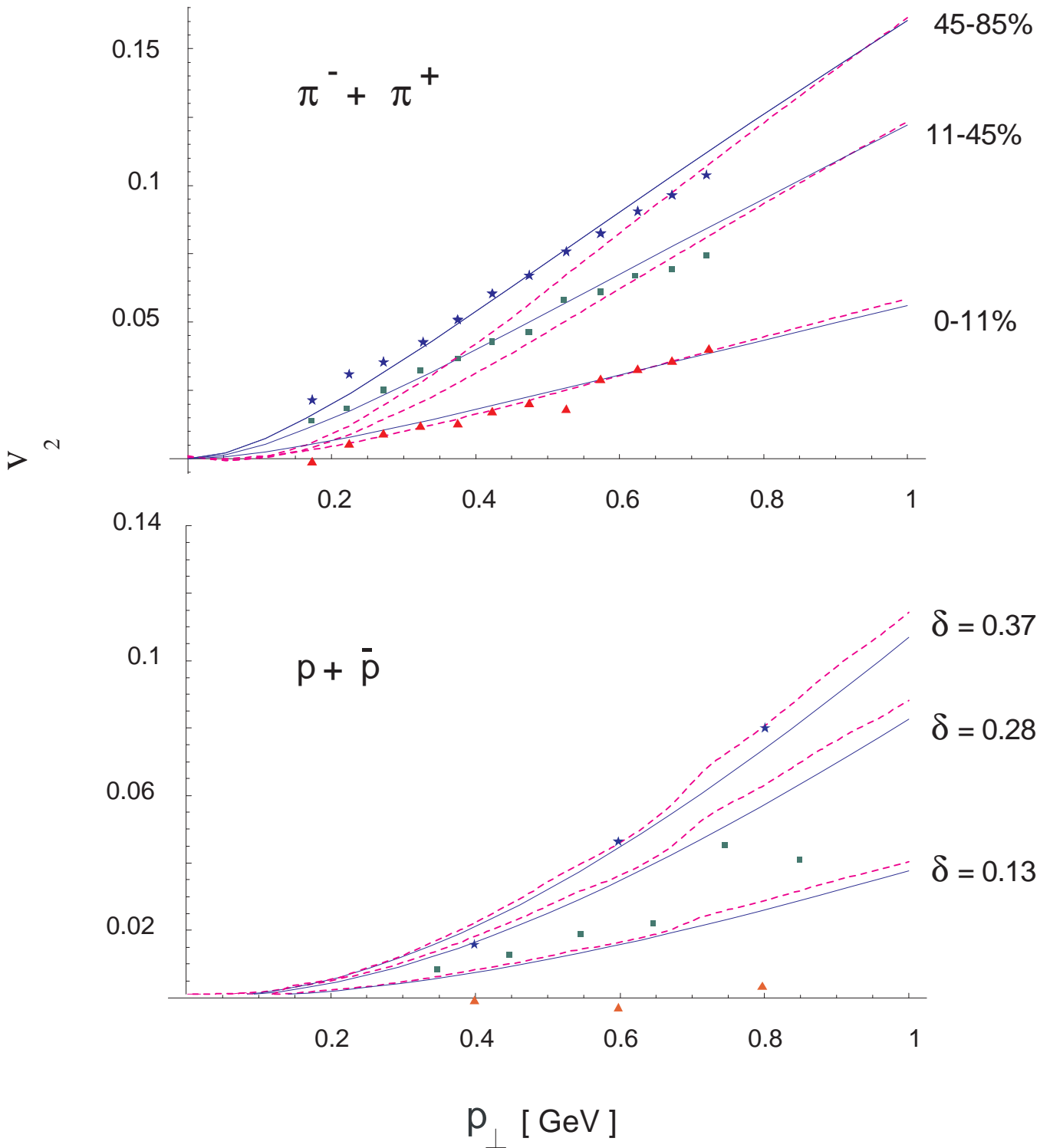


[data from STAR, M. Lisa, nucl-ex/0301005]

v_2 , PHENIX @ 130 GeV



v_2 , STAR @ 130 GeV



Questions to the model and beyond

1. Why should the simple $e^{-(E-\mu)/T}$ work? There is very little time to achieve thermal equilibrium in the gas of hadrons (early freeze-out)
2. Is this the property of **hadronization** ? (e.g. *fluctuating string tension*, A. Białas, Phys. Lett. B466 (1999) 301)
3. Significance of fluctuations; *data closer to hadron gas or constituent quark gas than to QGP* (A. Budzanowski, “The significance of M. Smoluchowski’s work in subatomic physics”, Acta Phys. Pol., in print)
4. Questions to prehistoric (*i.e.* pre-freeze-out) times: ***Was there quark-gluon plasma?*** If yes, why the transverse size does not grow with the collision energy?
5. How to construct a (microscopic) model for early stages such that the conditions at freeze-out which we use are reached?