

Fluktuacje v_2 – fizyka czy statystyka?*

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23.01.2009

*based on WB+P. Bożek+M. Rybczyński, *Fluctuating initial conditions in heavy-ion collisions from the Glauber approach*, Phys. Rev. C76 (2007) 054905 [arXiv:0706.4266]

Introduction

- Elliptic flow, measure:

$$v_2 = \frac{\langle p_x^2 \rangle - \langle p_y^2 \rangle}{\langle p_x^2 \rangle + \langle p_y^2 \rangle}$$

- Initial shape asymmetry:

$$\epsilon = \frac{\langle y^2 \rangle - \langle x^2 \rangle}{\langle x^2 \rangle + \langle y^2 \rangle}$$

- Hydro:

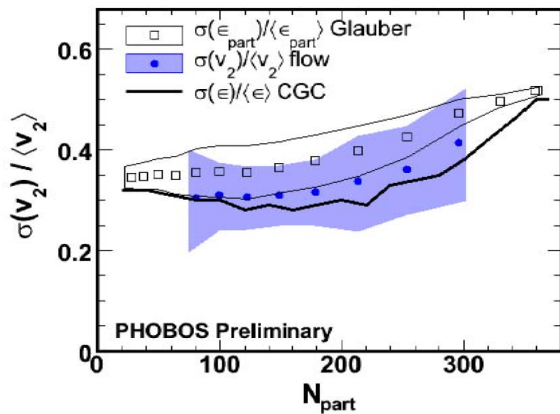
$$v_2 \sim \epsilon$$

(linearity of perturbation)

- Event-by-event fluctuations of v_2 measured (PHOBOS, STAR)

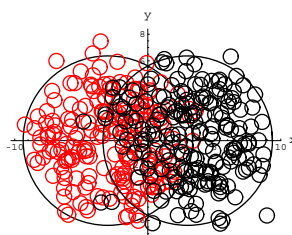
$$\Delta v_2 / v_2 \simeq \Delta \epsilon / \epsilon$$

Wosiek@WPCF'08

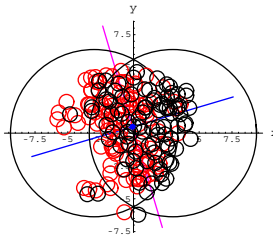


all end here!

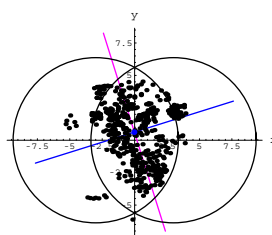
A typical gold-gold event



all nucleons



wounded nucleons



binary collisions

Sizable fluctuations of the principal axes

Aguiar+Kodama+Osada+Hama 2001, Miller+Snellings 2003, PHOBOS 2005,
Bhalerao+Blaizot+Borghini+Ollitrault 2005,
Andrade+Grassi+Hama+Kodama+Socolowski 2006, Voloshin 2006, ...

Toy model - n points on a unit circle

uniform distribution

$$f(\phi) = 1$$

fixed axes (standard)

$$\varepsilon = \frac{1}{n} \sum_{i=k}^n (y_k^2 - x_k^2) = \frac{1}{n} \sum_{k=1}^n \cos(2\phi_k)$$

Average over (infinitely many) events: $\langle \varepsilon \rangle = 0$

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variable axes (participant)

$$\varepsilon^* = \frac{1}{n} \sum_{k=1}^n \cos[2(\phi_k - \phi^*)]$$

$$Y = \frac{1}{n} \sum_{k=1}^n \cos(2\phi_k), \quad X = \frac{1}{n} \sum_{k=1}^n \sin(2\phi_k)$$

ϕ^* : quantity $\frac{1}{n} \sum_{k=1}^n \cos[2(\phi_k - \phi^*)]$ maximized in each event
 $\Rightarrow \quad \cos(2\phi^*) = Y/\sqrt{Y^2 + X^2}, \quad \sin(2\phi^*) = X/\sqrt{Y^2 + X^2}$

Central limit theorem

$$\varepsilon^* = \sqrt{\left(\frac{1}{n} \sum_{k=1}^n \cos(2\phi_k)\right)^2 + \left(\frac{1}{n} \sum_{k=1}^n \sin(2\phi_k)\right)^2} = \sqrt{X^2 + Y^2}$$

Unusual measure!

$$\langle X \rangle = \langle Y \rangle = 0, \quad \langle X^2 \rangle = \langle Y^2 \rangle = \frac{1}{2n}, \quad \langle XY \rangle = 0 \text{ (uncorrelated!)}$$

For large n the distribution of Y and X is Gaussian:

$$f(X, Y) = \frac{n}{\pi} \exp[-n(X^2 + Y^2)]$$

$$Y = q \cos \alpha, \quad X = q \sin \alpha, \quad f(q, \alpha) = \frac{n}{\pi} \exp[-nq^2]$$

Moments:

$$\begin{aligned} \int_0^{2\pi} d\alpha \int_0^\infty q dq f(q, \alpha) &= 1 \\ \int_0^{2\pi} d\alpha \int_0^\infty q dq q f(q, \alpha) &= \sqrt{\frac{\pi}{4n}} \equiv \langle \varepsilon^* \rangle \\ \int_0^{2\pi} d\alpha \int_0^\infty q dq q^2 f(q, \alpha) &= \frac{1}{n} \equiv \langle \varepsilon^{*2} \rangle \end{aligned}$$

$$\text{var}(\varepsilon^*) = \frac{1}{n} \left(1 - \frac{\pi}{4}\right)$$

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$$\text{var}(\varepsilon^*) = \frac{1}{n} \left(1 - \frac{\pi}{4}\right)$$

$$\frac{\sigma(\varepsilon^*)}{\langle \varepsilon^* \rangle} = \sqrt{\frac{4}{\pi} - 1} \simeq 0.52$$

Independent of n (for large n , i.e. in practice $n > 5$)

Toy model, $\epsilon \neq 0$

non-uniform distribution

$$f(\phi) = 1 + 2\epsilon \cos(2\phi)$$

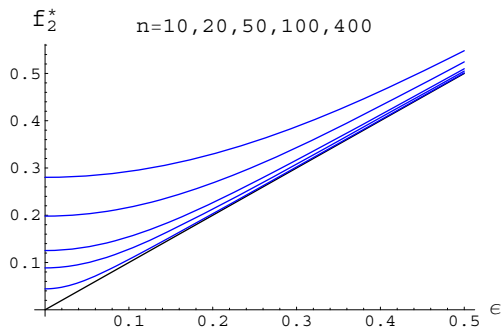
$$\epsilon^* = \frac{1 - 2\epsilon^2}{\sqrt{n\pi}} \sum_{j=0}^{\infty} (2\epsilon^2)^j \frac{\Gamma(j + \frac{1}{2})\Gamma(j + \frac{3}{2})}{j!^2} {}_1F_1\left(-\frac{1}{2}, j + 1; -\frac{n\epsilon^2}{1 - 2\epsilon^2}\right)$$

$$\langle \epsilon^{*2} \rangle = \frac{1 + (n - 1)\epsilon^2}{n}$$

limit of $\epsilon = 0$

$$\epsilon^*(\epsilon = 0) = \frac{\sqrt{\pi}}{2\sqrt{n}}, \quad \frac{\sigma(\epsilon^*)}{\epsilon^*} = \sqrt{\frac{4}{\pi} - 1}$$

ε^* as a function of ε (toy model)



ε^* in the general two-dimensional case

(under the assumption of no correlations of locations of sources)

$$\varepsilon^* = \frac{\sqrt{2}\sigma_Y^2}{I_{k,0}\sqrt{\pi}\sigma_X} \sum_{j=0}^{\infty} (2\delta\sigma_Y^2)^j \frac{\Gamma(j + \frac{1}{2}) \Gamma(j + \frac{3}{2}) {}_1F_1\left(-\frac{1}{2}; j + 1; -\frac{\bar{Y}^2}{2\sigma_Y^2}\right)}{j!^2}$$

$$\bar{Y} = I_{2,2}, \quad \sigma_Y^2 = \frac{1}{2n}(I_{4,0} - 2I_{2,2}^2 + I_{4,4}), \quad \sigma_X^2 = \frac{1}{2n}(I_{4,0} - I_{4,4}),$$

$$\delta = \frac{1}{2\sigma_Y^2} - \frac{1}{2\sigma_X^2}, \quad I_{k,l} = \int \rho d\rho f_l(\rho) \rho^k / \int \rho d\rho f_0(\rho)$$

at $b = 0$ very simple results (independent of A , energy, model, ...)

$$\varepsilon^* = \frac{\sqrt{\pi I_{4,0}}}{2I_{2,0}\sqrt{n}}, \quad \frac{\Delta\varepsilon^*}{\varepsilon^*} = \sqrt{\frac{4}{\pi} - 1}, \quad f_2^*(\rho) = \frac{1}{2} \sqrt{\frac{\pi}{nI_{2k,0}}} \rho^k f_0(\rho)$$

Glauber-like models tested, GLISSANDO

- wounded nucleons, $\sigma_w = 42$ mb, $d = 0.4$ fm

one goal:

compare various Glauber-like models

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 $\sigma_w = \sigma_{\text{bin}} = 42$ mb

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- hot spots: $\sigma_w = 42$ mb, $\sigma_{\text{bin}} = 0.5$ mb. When a rare binary collision occurs it produces on the average a large amount of the transverse energy = $14.5\% \times \sigma_w / \sigma_{\text{bin}}$

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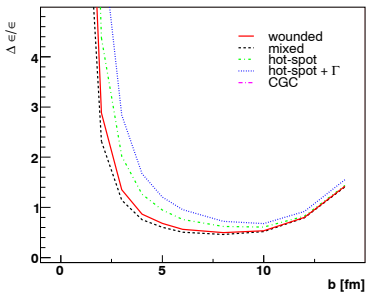
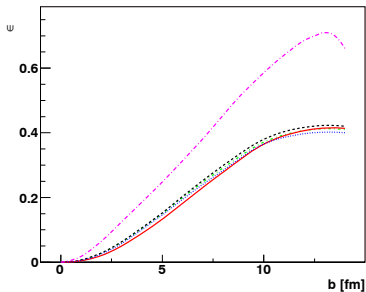
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- hot spots + Γ : Sources may deposit the transverse energy with a certain probability distribution. We superimpose the Γ distribution with $\kappa = 0.5$ over the distribution of sources,

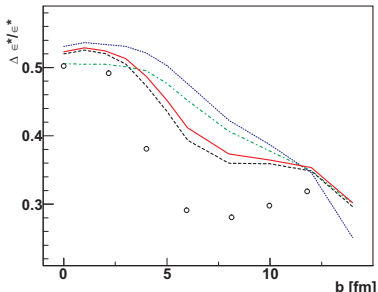
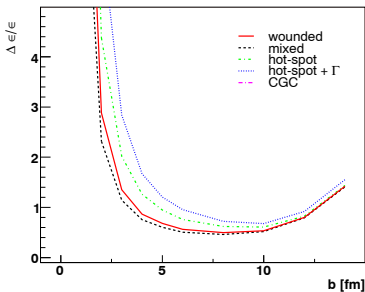
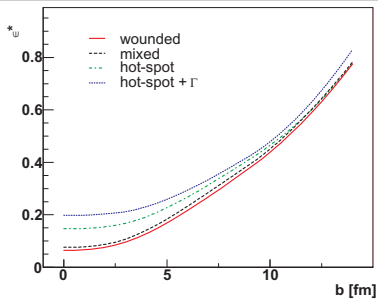
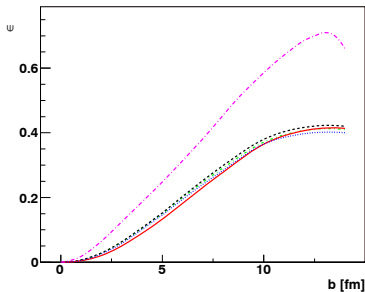
$$g(w, \kappa) = w^{\kappa-1} \kappa^\kappa \exp(-\kappa w) / \Gamma(\kappa),$$

where $\bar{w} = 1$ and $\text{var}(w) = 1/\kappa$

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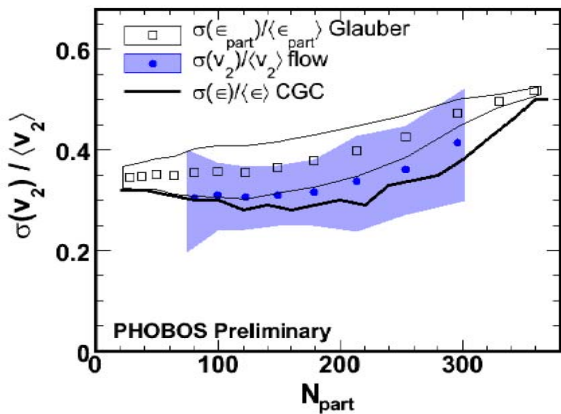


Event-by-event fluctuations of v_2

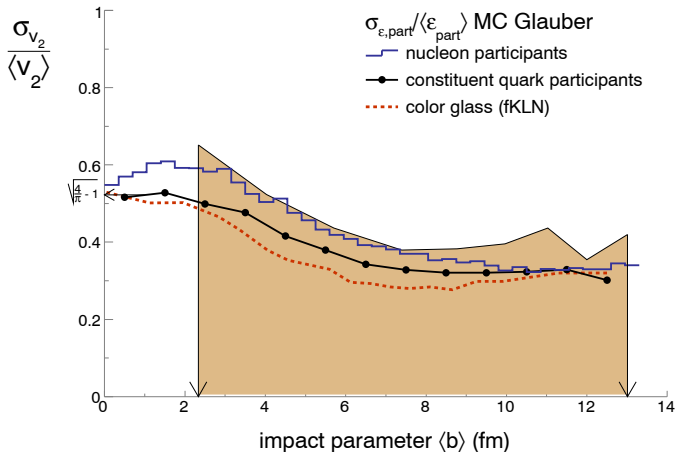
At low azimuthal asymmetry one expects on hydrodynamical grounds

$$\frac{\Delta v_2^*}{v_2^*} = \frac{\Delta \epsilon^*}{\epsilon^*}$$

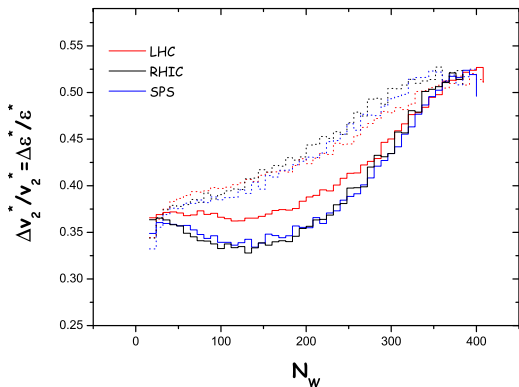
$$\frac{\Delta v_2^*}{v_2^*}(b=0) \simeq \frac{\Delta \epsilon^*}{\epsilon^*}(b=0) \simeq \sqrt{\frac{4}{\pi} - 1} \simeq 0.52$$



hydro: $\Delta\epsilon^*/\epsilon^* \simeq \Delta v_2^*/v_2^*$ (Au+Au, central) $\simeq 0.5$



[Sorensen, STAR]



from SPS (32mb), through RHIC (42mb) to LHC (63mb)

solid - wounded, dashed - hot-spot+ Γ

[Rybczyński, Epiphany'08]

Summary

- Analytic formulas explain why at $b = 0$ we have $\Delta\varepsilon^*/\varepsilon^* \simeq 0.52$, insensitive of the model used or the mass number of the colliding nuclei
- At $b > 0$ some sensitivity of $\Delta\varepsilon^*/\varepsilon^*$ on the model, small sensitivity on collision energy
- CGS [Drescher+Nara] has lower ε^* fluctuations than Glauber
- Comparison to data shows that the Glauber-like models leave little or no room for an increase of fluctuations during the (pre-hydro) evolution
- Not discussed: $\Delta v_4^*/v_4^* = 2\Delta v_2^*/v_2^*$